THE OBSERVER'S HANDBOOK FOR 1951

PUBLISHED BY

The Royal Istronomical Society of Canada

C. A. CHANT, EDITOR F. S. HOGG, Assistant Editor david dunlap observatory



FORTY-THIRD YEAR OF PUBLICATION

TORONTO 3 Willcocks Street Printed for the Society By the University of Toronto Press 1950

THE ROYAL ASTRONOMICAL SOCIETY OF CANADA

The Society was incorporated in 1890 as The Astronomical and Physical Society of Toronto, assuming its present name in 1903.

For many years the Toronto organization existed alone, but now the Society is national in extent, having active Centres in Montreal and Quebec, P.Q.; Ottawa, Toronto, Hamilton, London, Windsor, and Guelph, Ontario; Winnipeg, Man.; Saskatoon, Sask.; Edmonton, Alta.; Vancouver and Victoria, B.C. As well as nearly 1400 members of these Canadian Centres, there are nearly 500 members not attached to any Centre, mostly resident in other nations, while some 300 additional institutions or persons are on the regular mailing list of our publications. The Society publishes a monthly "Journal" and a yearly "Observer's Handbook". Single copies of the Journal are 50 cents, and of the Handbook, 40 cents.

Membership is open to anyone interested in astronomy. Annual dues, \$3.00; life membership, \$40.00. Publications are sent free to all members or may be subscribed for separately. Applications for membership or publications may be made to the Ceneral Secretary, 3 Willcocks St., Toronto.

CALENDAR

April Feb. Mar. Jan. SMTWTFS SMTWTF S SMTWTFS SMTWTFS 2 9 16 1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 26 27 7 8 14 15 2 3 3 1 2 3 4 5 6 7 1 5 12 19 4 4 5 6 6 7 10 11 12 13 14 17 18 19 20 21 24 25 26 27 28 8 9 10 10 8 9 15 16 22 23 11 12 13 14 15 18 19 20 21 22 16 17 11 13 17 22 23 24 18 20 21 22 23 24 25 26 27 28 29 30 31 28 29 30 31 25 26 27 28 29 30 July Aug. June May SMTWTFS SMTWTFS SMTWTFS SMTWTFS 3 4 4 6 7 10 11 12 13 14 19 19 20 21
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1951

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CONTENTS

											PAGE
Calendar	-	-	-	-	-	-	-	-	-	Cove	r p. ii
Preface	-	-	-	-	-	-	-	-	-	-	- 3
Anniversaries	and	Festiv	als	-	-	-	-	-	-	-	3
Symbols and	Abb	reviatio	ons	-	-	-	-	-	-	-	4
The Constella	tions	-	-	-	-	-	-	-	-	-	5
Miscellaneous	Astr	onomi	cal 🛛	Data	-	-	-	-	-	-	6
Ephemeris of	the	Sun	-	-	-	-	-	-	-	-	7
Solar and Side	ereal	Time	-	-	-	-	-	-	-	-	8
Julian Day C	alend	lar	-	-	-	-	-	-	-	-	8
Map of Stand	ard 7	Time Z	ones	s -	-	-	-	-	-	-	9
Times of Sunr	rise a	nd Sur	iset	-	-	-	-	-	-	-	10
Times of Beg	innin	g and	End	ling of	Twil	ight	-	-	-	-	17
Times of Moo	onrise	and M	1001	nset	-	-	-	-	-	-	18
The Planets f	for 1	951	-	-	-	-		-	-	-	24
Eclipses, 195	51	-	-	-	-	-	-	-	-	-	29
The Sky and A	Astro	nomica	l Pł	nenome	na M	onth	by M	onth	-	-	30
Phenomena o	f Jup	iter's S	Sate	llites	-	-	-	-	-	-	54
Lunar Occulta	ation	s, 1951	L	-	-	-	-	-	-	-	55
Meteors and I	Mete	orites	-	-	-	-	-	-	-	- 5	6, 80
Principal Elen	nents	of the	o So	lar Syst	tem	-	-	-	-	-	58
Satellites of th	he So	lar Sys	stem	ı - İ	-	-	-	-	-	-	59
Fields for Bri	ight `	Variabl	e S	tars	-	-	-	-	-	-	60
Representative	e Bri	ght V	arial	ble Sta	rs	-	-	-	-	-	61
Double and M	Multi	ole Sta	rs, v	with a	short	list	-	-	-	-	62
The Brightest	t Sta	rs, the	eir 🛛	magnitu	udes,	type	es, pr	oper	moti	ons,	
distances	and	radial	vel	ocities	-	-	-	-	-	-	64
Clusters and I	Nebu	lae:									
Star Clus	sters	-	-	-	-	-	-	-	-	-	72
Galactic	Neb	ulae	-	-	-	-	-	-	-	-	- 78
Extra-Ga	lactic	e Nebu	lae	-	-	-	-	-	-	-	74
Four Circular	: Star	Maps	5	-	-	-		-	-	-	75
Ephemeris for	the	Physic	al O	bservat	tion o	of the	Sun	-	-	-	79
Table of Prec	essio	n fór 5	60 Y	ears	-	-	-	-	-	Cover	p. iii

TABLES IN RECENT OBSERVER'S HANDBOOKS

Distance of the Stars—the Sun's Neighbours -	-	-	-	1941
Messier's List of Clusters and Nebulae	-	-	-	1 94 2
Meteorological Data: European and Asiatic -	-	-	-	1942
Canada and United States	-	-	-	1946
List of Air Navigation Stars	-	-	-	1947

PRINTED IN CANADA

The HANDBOOK for 1951 is the 43rd issue. During the past decade its circulation has increased from 1500 to 5500.

Four circular star maps 9 inches in diameter at a price of one cent each, and a set of four maps plotted on equatorial co-ordinates at a price of ten cents, are obtainable from the Director of University Extension, University of Toronto Toronto 5.

Celestial distances given herein are based on the standard value of 8".80 for the sun's parallax, not on the more recent value 8".790 determined by Sir

Harold Jones. Among the recent additions are: 1. Algol. Olin J. Eggen's epoch 2432520.6303 and period 2.86731525d., as published in the Astrophysical Journal, 1948.

2. Sun-spots. A table of solar rotation numbers for observers of sun-spots, and an ephemeris for physical observations of the sun.

Mr. Charles E. Apgar, of Westfield, New Jersey, died on August 17th at the age of eighty-five. For a number of years this enthusiastic amateur astronomer had prepared the tables of Jupiter's satellites for the HANDBOOK. The editors regret the death of this loyal friend.

Dr. F. S. Hogg, the Assistant Editor, as in recent years, assumed the responsibility of preparing this volume and to him the chief credit of its success is due; but sincere thanks are tendered to all those whose names are mentioned in the book, especially to Miss Ruth J. Northcott and Professor J. F. Heard. Our deep indebtedness to the British Nautical Almanac and the American Ephemeris is thankfully acknowledged.

C. A. CHANT.

David Dunlap Observatory, Richmond Hill, Ont., November, 1950.

ANNIVERSARIES AND FESTIVALS 1951

New Year's Day	1
EpiphanySat. Jan.	6
Septuagesima Sunday	21
Quinquagesima (Shrove	
Sunday)Feb.	4
Ash WednesdayFeb.	-7
St. David	1
St. PatrickSat. Mar.	17
Palm Sunday	18
Good Friday Mar.	23
Easter Sunday	25
St. George Mon. Apr.	23
Rogation SundayApr.	2 9
Ascension Day Thur. May	3
Pentecost (Whit Sunday) May	13
Trinity SundayMay	20
Empire Day (Victoria	
Day)	24
Corpus Christi	24
Birthday of the Queen Mother,	
Mary (1867)	30
St. John Baptist (Mid-	
Summer Dav)Sun. June	24

Dominion DaySun. Birthday of Queen	July	1
Elizabeth (1900)Sat.	Aug.	4
Labour Day	Sept.	3
St. Michael (Michaelmas		
DaySat.	Sept.	29
Hebrew New Year (Rosh	•	
Hashanah)Mon.	Oct.	1
All Saints' Day Thur.	Nov.	1
Remembrance DaySun.	Nov.	11
St. AndrewFri.	Nov.	30
First Sunday in Advent	Dec.	2
Accession of King		
George VI (1936)Tues.	Dec.	11
Birthday of King		
George VI (1895)Fri.	Dec.	14
Christmas DayTues.	Dec.	25

Thanksgiving Day, Date set by Proclamation.

SIGNS OF THE ZODIAC

'''Aries 0°	Ω Leo 120°	オ Sagittarius240 [™]
∀ Taurus 30°	110° Wirgo 150°	o Capricornus 270°
¤ Gemini	\simeq Libra	a Aquarius
6 Cancer	\mathfrak{m} Scorpio $\ldots 210^{\circ}$) (Pisces

SUN, MOON AND PLANETS

Ó	The Sun.	Q	The Moon generally.	24	Jupiter.
•	New Moon.	ĝ	Mercury.	Þ	Saturn.
٢	Full Moon.	ę	Venus.	ô	or H Uranus.
D	First Quarter	Ð	Earth.	Ψ	Neptune.
đ	Last Quarter.	ð	Mars.	E	Pluto

ASPECTS AND ABBREVIATIONS

σ' Conjunction, or having the same Longitude or Right Ascension
♂ Opposition, or differing 180° in Longitude or Right Ascension
□ Quadrature, or differing 90° in Longitude or Right Ascension.
□ A Ascending Node; ♡ Descending Node.
a or A. R., Right Ascension; δ Declination.
h, m, s, Hours, Minutes, Seconds of Time.
*' ", Degrees, Minutes, Seconds of Arc.

THE GREEK ALPHABET

A, a,	Alpha.	Ι,ι,	Iota.	Ρ,ρ,	Rho.
$\mathbf{B}, \boldsymbol{\beta},$	Beta.	Κ, κ,	Kappa.	Σ,σ,ς,	Sigma.
Γ,γ,	Gamma.	Λ, λ,	Lambda.	Τ, τ,	Tau.
Δ,δ,	Delta.	Μ,μ,	Mu.	Υ, ν,	Upsilon.
Εε	Epsilon.	Ν, ν,	Nu.	Φ, φ,	Phi.
Ζ,ζ,	Zeta.	Ξ,ξ,	Xi.	Χ, χ,	Chi.
Η, η,	Eta.	0,0,	Omi cro n.	Ψ,ψ,	Psi.
θ,θ,ϑ,	Theta.	Π,π,	Pi.	Ω,ω,	Omega.

THE CONFIGURATIONS OF JUPITER'S SATELLITES

In the Configurations of Jupiter's Satellites (pages 31, 33, etc.), O represents the disc of the planet, d signifies that the satellite is on the disc, * signifies that the satellite is behind the disc or in the shadow. Configurations are for an inverting telescope.

THE CONSTELLATIONS

LATIN AND ENGLISH NAMES WITH ABBREVIATIONS

Andromeda,		Leo, <i>Lion</i> Leo	Leon
(Chained Maiden) And	Andr	Leo Minor, Lesser Lion LMi	LMin
Antlia, Air PumpAnt	Antl	Lepus, HareLep	Leps
Apus, Bird of Paradise Aps	Apus	Libra, ScalesLib	Libr
Aquarius, Water-bearer Agr	Agar	Lupus, WolfLup	Lupi
Aquila, EagleAql	Aqil	Lynx, LynxLyn	Lync
Ara, AltarAra	Arae	Lvra, LyreLvr	Lyra
Aries, Ram Ari	Arie	Mensa, Table (Mountain) Men	Mens
Auriga. (Charioteer)Aur	Auri	Microscopium.	
Bootes, (Herdsman)Boo	Boot	Microscope Mic	Micr
Caelum. Chisel	Cael	Monoceros, Unicorn Mon	Mono
Camelopardalis, Giraffe., Cam	Caml	Musca, FlyMus	Musc
Cancer. CrabCnc	Canc	Norma, Square,	Norm
Canes Venatici.	ouno	Octans. Octant Oct	Octn
Hunting DogsCVn	CVen	Ophiuchus.	ocen
Canis Major, Greater Dog CMa	CMai	Serbent-bearer Oph	Onhi
Canis Minor Lesser Dog CMi	CMin	Orion (Hunter) Ori	Orio
Capricornus Sea-goat Cap	Capr	Pavo Pencock Pav	Pavo
Carina Keel Car	Cari	Pegasus (Winged Horse) Peg	Pege
Cassioneia	Carr	Perseus (Chambian) Per	Pers
(Lady in Chair) Cas	Case	Phoenix Phaenix Phe	Phoe
Centaurus Centaur Cen	Cent	Pictor Painter Pic	Pict
Cepheus (King) Cep	Cenh	Piscos Fishes Psc	Pice
Cetus Whale Cet	Ceti	Piecie Australie	1 150
Chamaeleon Chamaeleon Cha	Cham	Southarn Fish De A	PecA
Circinus Compassas Cir	Circ	Puppis Pood Pup	Pupp
Columba Dava Col	Colm	Purio Compass Pur	Duni
Como Baranicas	Com	Potioulum Nat Dot	r yxi Doti
Baranica's Hair Com	Como	Sogitto Annon Sag	Sete
Company Australia	Coma	Sagitta, Arrow	Sgte
Southern Crown Cr	ConA	Sagittarius, Archer	Sgtr
Southern CrownCrA	COFA	Scorpius, ScorpionScr	Scor
Northann Creans,	CarD	Sculptor, Sculptor Sci	Scul
Northern Crown Crb	Corb	Scutum, SnielaSct	Scut
Corvus, CrowCrv	Corv	Serpens, SerpentSer	Serp
Crater, CupCrt	Crat	Sextans, SextantSex	Sext
Crux, (Soumern) Cross. Cru	Cruc	Taurus, Buu	laur
Cygnus, SwanCyg	Cygn	Telescopium, TelescopeTel	lele
Delphinus, DolphinDel	Diph	Triangulum, Triangle Iri	Tria
Dorado, SwordfishDor	Dora	Triangulum Australe,	m 4
Draco, DragonDra	Drac	Southern Triangle TrA	TrAu
Equuleus, Little HorseEqu	Equi	Tucana, ToucanTuc	Tucn
Eridanus, River Eridanus. Eri	Erid	Ursa Major, Greater Bear. UMa	UMaj
Fornax, FurnaceFor	Forn	Ursa Minor, Lesser Bear. UMi	UMin
Gemini, Twins	Gemi	Vela, SailsVel	Velr
Grus, CraneGru	Grus	Virgo, VirginVir	Virg
Hercules,		Volans, Flying FishVol	Voln
(Kneeling Giant) Her	Herc	Vulpecula, FoxVul	Vulp
Horologium, ClockHor	Horo		
Hydra, Water-snakeHya	Hyda	The 4-letter abbreviations	are in-
Hydrus, Sea-serpentHyi	Hydi	tended to be used in cases w	vhere a
Indus, IndianInd	Indi	maximum saving of space	is not
Lacerta, <i>Lizard</i> Lac	Lacr	necessary.	

UNITS OF LENGTH 1 Angstrom unit = 10^{-8} cm. 1 micron = 10-4 cm. 1 meter $= 10^{\circ}$ cm. = 3.28084 feet 1 kilometer = 10⁵ cm. = 0.62137 miles 1 mile = 1.60935 ×10⁵ cm. = 1.60935 km. 1 astronomical unit = 1.49504 × 1013 cm. = 92,897,416 miles 1 light year = 9.463 × 10¹⁷ cm. = 5.880 × 10¹² miles = 0.3069 parsecs 1 parsec $= 30.84 \times 10^{17}$ cm. $= 19.16 \times 10^{12}$ miles = 3.259 l.y. 1 megaparsec = 30.84×10^{28} cm. = 19.16×10^{18} miles = 3.259×10^{6} l.y. UNITS OF TIME Sidereal day = 23h 56m 04.09s of mean solar time Mean solar day = 24h 03m 56.56s of sidereal time Synodical month = $29d \ 12h \ 44m$; sidereal month = $27d \ 07h \ 43m$ Tropical year (ordinary) = 365d 05h 48m 46s Sidereal year = 365d 06h 09m 10s Eclipse year =346d 14h 53m THE EARTH Equatorial radius, a = 3963.35 miles; flattening, c = (a-b)/a = 1/297.0Polar radius, b = 3950.01 miles 1° of latitude = $69.057 - 0.349 \cos 2\phi$ miles (at latitude ϕ) 1° of longitude = 69.232 cos ϕ -0.0584 cos 3 ϕ miles Mass of earth = 6.6×10^{21} tons; velocity of escape from $\bigoplus = 6.94$ miles/sec. EARTH'S ORBITAL MOTION Solar parallax = 8.''80; constant of aberration = 20.''47Annual general precession = 50."26; obliquity of ecliptic = 23° 26' 50" (1939) Orbital velocity = 18.5 miles/sec.; parabolic velocity at \bigoplus = 26.2 miles/sec SOLAR MOTION Solar apex, R.A. 18h 04m; Dec. + 31° Solar velocity = 12.2 miles/sec. THE GALACTIC SYSTEM North pole of galactic plane R.A. 12h 40m, Dec. + 28° (1900) Centre. 325° galactic longitude, = R.A. 17h 24m, Dec. -30° Distance to centre = 10,000 parsecs; diameter = 30,000 parsecs. Rotational velocity (at sun) = 262 km./sec. Rotational period (at sun) = $2.2 \times 10^{\circ}$ years Mass = 2×10^{11} solar masses EXTRAGALACTIC NEBULAE Red shift =+530 km./sec./megaparsec=+101 miles /sec./million l.y. **RADIATION CONSTANTS** Velocity of light = 299,774 km./sec. = 186,271 miles/sec. Solar constant = 1.93 gram calories/square cm./minute Light ratio for one magnitude = 2.512; log ratio = 0.4000 Radiation from a star of zero apparent magnitude = 3×10^{-6} meter candles Total energy emitted by a star of zero absolute magnitude = 5×10^{25} horsepower MISCELLANEOUS Constant of gravitation, $G = 6.670 \times 10^{-8}$ c.g.s. units Mass of the electron, $m = 9.035 \times 10^{-28}$ gm.; mass of the proton = 1.662×10^{-14} gm Planck's constant, $h = 6.55 \times 10^{-27}$ erg. sec. Loschmidt's number = 2.705×10^{19} molecules/cu. cm. of gas at N.T.P. Absolute temperature = T° K = T° C + 273° = 5/9 (T° F + 459°) 1 radian = 57°.2958 $\pi = 3.141,592,653,6$ = 3437'.75 No. of square degrees in the sky = 206.265''~ 41,253

1951 EPHEMERIS OF THE SUN AT 0h GREENWICH CIVIL TIME

Date 1951	Apparent R.A.	t Corr. to Apparent Sun-dial Dec.		Date 1951	Apparent R.A.	Corr. to Sun-dial	Apparent Dec.
Jan. 1 4 7 10 13 16 19 22 25 28 31	h m s 18 42 27 18 55 41 19 08 52 19 21 59 19 35 01 19 47 57 20 00 47 20 13 31 20 26 07 20 38 37 20 50 59	$\begin{array}{c} m & s \\ +03 & 06 \\ +04 & 30 \\ +05 & 52 \\ +07 & 09 \\ +08 & 21 \\ +09 & 28 \\ +10 & 28 \\ +11 & 22 \\ +12 & 09 \\ +12 & 49 \\ +13 & 21 \end{array}$	$ \begin{array}{c} \circ & , \\ -23 & 05.4 \\ -22 & 49.9 \\ -22 & 30.3 \\ -22 & 06.7 \\ -21 & 07.9 \\ -20 & 33.0 \\ -19 & 54.6 \\ -19 & 12.9 \\ -18 & 28.0 \\ -17 & 40.1 \end{array} $	Jul. 3 6 9 12 15 18 21 24 27 30	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c} \circ & i \\ +23 & 03.0 \\ +22 & 48.0 \\ +22 & 29.3 \\ +21 & 07.2 \\ +21 & 14.7 \\ +21 & 12.9 \\ +20 & 40.8 \\ +20 & 05.6 \\ +19 & 27.4 \\ +18 & 46.2 \end{array}$
Feb. 3 6 9 12 15 18 21 24 27	21 03 14 21 15 22 21 27 23 21 39 16 21 51 03 22 02 42 22 14 16 22 25 43 22 37 05	$\begin{array}{r} +13 & 47 \\ +14 & 05 \\ +14 & 16 \\ +14 & 20 \\ +14 & 17 \\ +14 & 07 \\ +13 & 50 \\ +13 & 28 \\ +13 & 00 \end{array}$	$\begin{array}{c} -16 & 49.3 \\ -15 & 55.9 \\ -15 & 00.0 \\ -14 & 01.9 \\ -13 & 01.6 \\ -11 & 59.5 \\ -10 & 55.7 \\ -09 & 50.4 \\ -08 & 43.7 \end{array}$	Aug. 2 5 8 11 14 17 20 23 26 29	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{r} +06 \ 14 \\ +06 \ 01 \\ +05 \ 42 \\ +05 \ 18 \\ +04 \ 49 \\ +04 \ 15 \\ +03 \ 36 \\ +02 \ 52 \\ +02 \ 05 \\ +01 \ 14 \end{array}$	$\begin{array}{r} +18 & 02.3 \\ +17 & 15.7 \\ +16 & 26.6 \\ +15 & 35.1 \\ +14 & 41.4 \\ +13 & 45.6 \\ +12 & 47.8 \\ +11 & 48.2 \\ +10 & 46.9 \\ +09 & 44.1 \end{array}$
Mar. 2 5 8 11 14 17 20 23 26 29	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c} +12 & 27 \\ +11 & 50 \\ +11 & 09 \\ +10 & 24 \\ +09 & 37 \\ +08 & 46 \\ +07 & 54 \\ +07 & 00 \\ +06 & 05 \\ +05 & 10 \end{array}$	$\begin{array}{c} -07 & 35.9 \\ -06 & 27.0 \\ -05 & 17.3 \\ -04 & 07.0 \\ -02 & 56.2 \\ -01 & 45.2 \\ -00 & 34.0 \\ +00 & 37.1 \\ +01 & 48.0 \\ +02 & 58.5 \end{array}$	Sep. 1 4 7 10 13 16 19 22 25 28	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{c} +08 & 39.8 \\ +07 & 34.3 \\ +06 & 27.7 \\ +05 & 20.2 \\ +04 & 11.8 \\ +03 & 02.8 \\ +01 & 53.3 \\ +00 & 43.5 \\ -00 & 26.7 \\ -01 & 36.8 \end{array}$
Apr. 1 4 7 10 13 16 19 22 25 28	00 38 27 00 49 23 01 00 20 01 11 19 01 22 21 01 33 25 01 44 32 01 55 43 02 06 57 02 18 16	$\begin{array}{r} +04 \ 16 \\ +03 \ 22 \\ +02 \ 30 \\ +01 \ 39 \\ +00 \ 51 \\ +00 \ 06 \\ -00 \ 37 \\ -01 \ 16 \\ -01 \ 51 \\ -02 \ 22 \end{array}$	$\begin{array}{c} +04 & 08.5 \\ +05 & 17.9 \\ +06 & 26.4 \\ +07 & 33.9 \\ +08 & 40.3 \\ +09 & 45.4 \\ +10 & 48.9 \\ +11 & 50.9 \\ +12 & 51.1 \\ +13 & 49.5 \end{array}$	Oct. 1 4 7 10 13 16 19 22 25 28 31	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{r} -02 \ 46.9 \\ -03 \ 56.8 \\ -05 \ 06.2 \\ -06 \ 15.0 \\ -07 \ 23.1 \\ -08 \ 30.3 \\ -09 \ 36.3 \\ -10 \ 41.2 \\ -11 \ 44.6 \\ -12 \ 46.4 \\ -13 \ 46.5 \end{array}$
May 1 4 7 10 13 16 19 22 25 28 31	02 29 39 02 41 08 02 52 41 03 04 20 03 16 03 03 27 52 03 39 45 03 51 44 04 03 47 04 15 55 04 28 07	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{c} +14 \ 45.8 \\ +15 \ 39.9 \\ +16 \ 31.7 \\ +17 \ 21.0 \\ +18 \ 07.7 \\ +18 \ 51.7 \\ +19 \ 32.8 \\ +20 \ 10.8 \\ +20 \ 45.8 \\ +21 \ 17.6 \\ +21 \ 46.1 \end{array}$	Nov. 3 6 9 12 15 18 21 24 27 30	14 29 24 14 41 16 14 53 14 15 05 20 15 17 33 15 29 54 15 42 23 15 54 59 16 07 42 16 20 32	$\begin{array}{c} -16 & 23 \\ -16 & 21 \\ -16 & 13 \\ -15 & 56 \\ -15 & 33 \\ -15 & 01 \\ -14 & 22 \\ -13 & 36 \\ -12 & 42 \\ -11 & 42 \end{array}$	$\begin{array}{c} -14 & 44.6 \\ -15 & 40.5 \\ -16 & 34.1 \\ -17 & 25.1 \\ -18 & 13.5 \\ -18 & 58.9 \\ -19 & 41.4 \\ -20 & 20.6 \\ -20 & 56.5 \\ -21 & 28.8 \end{array}$
Jun. 3 6 9 12 15 18 21 24 27 30	04 40 23 04 52 43 05 05 06 05 17 31 05 29 58 05 42 26 05 54 54 06 07 23 06 19 51 06 32 17	$\begin{array}{r} -02 & 11 \\ -01 & 41 \\ -01 & 07 \\ -00 & 32 \\ +00 & 05 \\ +00 & 43 \\ +01 & 22 \\ +02 & 01 \\ +02 & 39 \\ +03 & 16 \end{array}$	$\begin{array}{r} +22 & 11.2 \\ +22 & 32.8 \\ +22 & 50.9 \\ +23 & 05.3 \\ +23 & 16.1 \\ +23 & 23.2 \\ +23 & 26.6 \\ +23 & 26.2 \\ +23 & 26.2 \\ +23 & 14.4 \end{array}$	Dec. 3 6 9 12 15 18 21 24 27 30	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c} -10 & 35 \\ -09 & 23 \\ -08 & 07 \\ -06 & 46 \\ -05 & 52 \\ -03 & 55 \\ -02 & 26 \\ -00 & 56 \\ +00 & 33 \\ +02 & 02 \end{array}$	$\begin{array}{rrrrr} -21 & 57.5 \\ -22 & 22.4 \\ -22 & 43.3 \\ -23 & 00.2 \\ -23 & 13.0 \\ -23 & 21.7 \\ -23 & 26.2 \\ -23 & 26.4 \\ -23 & 22.4 \\ -23 & 14.2 \end{array}$

SOLAR AND SIDEREAL TIME

In practical astronomy three different kinds of time are used, while in ordinary life we use a fourth.

1. Apparent Time—By apparent noon is meant the moment when the sun is on the meridian, and apparent time is measured by the distance in degrees that the sun is east or west of the meridian. Apparent time is given by the sun-dial.

2. Mean Time—The interval between apparent noon on two successive days is not constant, and a clock cannot be constructed to keep apparent time. For this reason mean time is used. The length of a mean day is the average of all the apparent days throughout the year. The real sun moves about the ecliptic in one year; an imaginary mean sun is considered as moving uniformly around the celestial equator in one year. The difference between the times that the real sun and the mean sun cross the meridian is the equation of time. Or, in general, Apparent Time—Mean Time = Equation of Time. This is the same as Correction to Sun-dial on page 7, with the sign reversed.

3. Sidereal Time—This is time as determined from the stars. It is sidereal noon when the Vernal Equinox or First of Aries is on the meridian. In accurate time-keeping the moment when a star is on the meridian is observed and the corresponding mean time is then computed with the assistance of the Nautical Almanac. When a telescope is mounted equatorially the position of a body in the sky is located by means of the sidereal time.

4. Standard Time—In everyday life we use still another kind of time. A moment's thought will show that in general two places will not have the same mean time; indeed, difference in longitude between two places is determined from their difference in time. But in travelling it is very inconvenient to have the time varying from station to station. For the purpose of facilitating transportation the system of *Standard Time* was introduced in 1883. Within a certain belt approximately 15° wide, all the clocks show the same time, and in passing from one belt to the next the hands of the clock are moved forward or backward one hour.

In Canada we have seven standard time belts, as follows;—Newfoundland Time, 3h. 30m. slower than Greenwich; 60th meridian or Atlantic Time, 4h.; 75th meridian or Eastern Time, 5h.; 90th meridian or Central Time, 6h.; 105th meridian or Mountain Time, 7h.; 120th meridian or Pacific Time, 8h.; and 135th meridian or Yukon Time, 9h. slower than Greenwich.

The boundaries of the time belts are shown on the map on page 9.

Daylight Saving Time is the standard time of the next zone eastward. It is adopted in many places between certain specified dates during the summer.



Revisions: Newfoundland Time is 3h. 30m. slower than Greenwich Time. The "panhandle" region of Alaska, containing such towns as Juneau and Skagway, is on 120th meridian (Pacific) Time, instead of Yukon Time.

JULIAN CALENDAR, 1951 J.D. 2,430,000 plus the following Jan. 1 Feb. 1 3648 May 1 3768 Sep. 1 Oct. 1 3891 3799 3829 3679 Jun. 1 3921 Mar. 1 3707 Jul. 1 3829 Aug. 1 3860 Nov. 1 3952 Nov. 1 3952 Dec. 1 3982 3738 Apr. 1 The Julian Day commences at noon. Thus J.D. 2,433,648 = Jan. 1.5 G.C.T.

TIMES OF SUNRISE AND SUNSET

In the tables on pages 11 to 16 are given the times of sunrise and sunset for places in latitudes 32° , 36° , 40° , 44° , 46° , 48° , 50° , and 52° . The times are given in Local Mean Time, and in the table below are given corrections to change from Local Mean to Standard Time for the cities and towns named.

The time of sunrise and sunset at a given place, in local mean time, varies from day to day, and depends principally upon the declination of the sun. Variations in the equation of time, the apparent diameter of the sun and atmospheric refraction at the points of sunrise and sunset also affect the final result. These quantities, as well as the solar declination, do not have precisely the same values on corresponding days from year to year, and so the table gives only approximately average values. The times are for the rising and setting of the upper limb of the sun, and are corrected for refraction. It must also be remembered that these times are computed for the sea horizon, which is only approximately realised on land surfaces.

The Standard Times for Any Station

In order to find the time of sunrise and sunset for any place on any day, first from the list below find the approximate latitude of the place and the correction, in minutes, which follows the name. Then find in the monthly table the local time of sunrise and sunset for the proper latitude, on the desired day, and apply the correction to get the Standard Time.

CANADIAN CITIES AND TOWNS						AMERICAN CITIES			
	Lat.	Cor.		Lat.	Cor.		Lat.	Cor.	
Belleville Brandon Brantford Calgary Charlottetown Chatham Cornwall Dawson Edmonton Fort William Fredericton Galt Glace Bay Guelph Halifax Hamilton Hull Kingston Kitchener London Medicine Hat Montcal Monse Jaw Niagara Falls North Bay Oshawa Ottawa	$\begin{array}{r} 44\\ 450\\ 43\\ 51\\ 42\\ 45\\ 44\\ 54\\ 43\\ 64\\ 43\\ 45\\ 44\\ 43\\ 45\\ 43\\ 45\\ 44\\ 43\\ 50\\ 64\\ 45\\ 50\\ 34\\ 64\\ 45\\ 50\\ 34\\ 64\\ 45\\ 50\\ 50\\ 50\\ 50\\ 50\\ 50\\ 50\\ 50\\ 50\\ 5$	$\begin{array}{c} +99\\ ++216\\ ++139\\ ++139\\ ++134\\ +-018\\ ++134\\ ++193\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ ++193\\118\\ +-118\\ +-118\\18$	Peterborough Port Arthur Prince Albert Prince Rupert Quebec Regina St. Catharines St. Hyacinthe Saint John, N.B. St. John's, Nfld. St. Thomas Sarnia Saskatoon Sault Ste. Marie Sharbrooke Stratford Sudbury Sydney Timmins Toronto Three Rivers Trail Truro Vancouver Victoria Windsor Winnipeg Woodstock	$\begin{array}{c} 44\\ 48\\ 53\\ 54\\ 7\\ 50\\ 43\\ 46\\ 45\\ 43\\ 43\\ 22\\ 47\\ 45\\ 43\\ 47\\ 46\\ 48\\ 44\\ 46\\ 9\\ 45\\ 48\\ 42\\ 9\\ 45\\ 48\\ 42\\ 50\\ 53\\ 32\\ 68\\ 68\\ 68\\ 68\\ 68\\ 68\\ 68\\ 68\\ 68\\ 68$	$\begin{array}{c} +13\\ +57\\ +031\\ +152\\ -107\\ +24\\ +102\\ +307\\ +37\\ -102\\ +201\\ +20$	Atlanta Baltimore Birmingham Boston Buffalo Chicago Cincinnati Cleveland Dallas Denver Detroit Fairbanks Indianapolis Juneau Kansas City Los Angeles Louisville Memphis Milwaukee Minneapolis New Orleans New York Omaha Philadelphia Pittsburgh Portland St. Louis San Francisco Seattle	$\begin{array}{c} 34\\ 39\\ 42\\ 43\\ 42\\ 39\\ 42\\ 33\\ 40\\ 42\\ 65\\ 58\\ 39\\ 43\\ 8\\ 5\\ 30\\ 41\\ 40\\ 46\\ 39\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\$	$\begin{array}{c} +37\\ +106\\ -118\\ +115\\ +115\\ +27\\ 002\\ +105\\ +27\\ 002\\ +105\\ -158\\ +27\\ 009\\ +210\\ -158\\ +201\\ -170\\ 009\\ +201\\ +$	
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Example—Find the time of sunrise at Owen Sound, on February 12.

In the above list Owen Sound is under " 45° ", and the correction is + 24 min. On page 11 the time of sunrise on February 12 for latitude 45° is 7.07; add 24 min. and we get 7.31 (Eastern Standard Time).

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	BEGINNING	OF	MORNING A	AND	ENDING OI	F EVENING	TWILIGH
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	Latitude 35°	Latitude 40°	Latitude 45°	Latitude 50°	Latitude 52°
	Morn. Eve.	Morn. Eve.	Morn. Eve.	Morn. Eve.	Morn. Eve.
Jan. 1 11 21 31 Feb. 10	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
20 Mar. 2 12 22 Apr. 1	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
11 21 May 1 11 21	4 07 7 57 3 51 8 07 3 37 8 19 3 23 8 30 3 12 8 41	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
31 June 10 20 30 July 10	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	2 36 9 20 2 29 9 30 2 27 9 35 2 31 9 35 2 39 9 30	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	0 23 11 42 	
20 30 Aug. 9 19 29	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$
Sept. 8 18 28 Oct. 8	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
Nov. 28 Nov. 7 17 27 Dec. 7	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
17 27 Jan. 1	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$

The above table gives the local mean time of the beginning of morning twilight, and of the ending of evening twilight, for various latitudes. To obtain the corresponding standard time, the method used is the same as for correcting the sunrise and sunset tables, as described on page 10. The entry — in the above table indicates that at such dates and latitudes, twilight lasts all night. This table, taken from the American Expenseries, is computed for *astronomical* twilight, i.e. for the time at which the sun is 108' from the zenith (or 18° below the horizon).

FIMES OF MOONRISE AND	MOONSET, 1951.	(Local Mean	Time)
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Date	Latitu Mo Rise	de 35° oon Set	Latitu Mo Rise	de 40° oon Set	Latitu Mo Rise	de 45° oon Set	Latitu Mo Rise	de 50° oon Set	Latitu Mo Rise	de 52° Don Set
Jan. 1 (2 3 4 5	h m 00 10 01 17 02 27 03 42 04 58	h m 11 51 12 21 12 57 13 40 14 33	h m 00 14 01 24 02 39 03 58 05 17	h m 11 46 12 12 12 44 13 23 14 14	h m 00 17 01 32 02 52 04 16 05 38	h m 11 38 12 02 12 29 13 03 13 51	h m 00 23 01 41 03 10 04 39 06 07	h m 11 34 11 50 12 10 12 39 13 22	h m 00 24 01 48 03 17 04 50 06 21	h m 11 31 11 44 12 02 12 28 13 07
6 7 0 8 9 10	06 10 07 14 08 07 08 50 09 24	$\begin{array}{cccc} 15 & 37 \\ 16 & 50 \\ 18 & 06 \\ 19 & 20 \\ 20 & 29 \end{array}$	06 30 07 33 08 23 09 02 09 32	$\begin{array}{cccc} 15 & 17 \\ 16 & 31 \\ 17 & 50 \\ 19 & 08 \\ 20 & 22 \end{array}$	06 55 07 56 08 43 09 17 09 43	$\begin{array}{rrrr} 14 & 53 \\ 16 & 09 \\ 17 & 32 \\ 18 & 55 \\ 20 & 14 \end{array}$	07 25 08 25 09 06 09 34 09 55	14 23 15 40 17 09 18 39 20 04	07 40 08 39 09 18 09 43 10 01	14 07 15 27 16 58 18 31 19 59
11 12 13 14 15	09 53 10 19 10 44 11 08 11 34	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	09 58 10 20 10 41 11 03 11 25	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 10 & 04 \\ 10 & 22 \\ 10 & 39 \\ 10 & 56 \\ 11 & 14 \end{array}$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 10 & 11 \\ 10 & 24 \\ 10 & 36 \\ 10 & 48 \\ 11 & 02 \end{array}$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$
16 17 18 19 20	$\begin{array}{rrrrr} 12 & 03 \\ 12 & 36 \\ 13 & 14 \\ 13 & 58 \\ 14 & 50 \end{array}$	01 35 02 35 03 34 04 32 05 26	$ \begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{cccc} 01 & 46 \\ 02 & 49 \\ 03 & 52 \\ 04 & 51 \\ 05 & 45 \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 02 & 00 \\ 03 & 07 \\ 04 & 12 \\ 05 & 14 \\ 06 & 09 \end{array}$	11 18 11 40 12 08 12 46 13 37	$\begin{array}{cccc} 02 & 16 \\ 03 & 28 \\ 04 & 38 \\ 05 & 43 \\ 06 & 40 \end{array}$	11 11 11 29 11 55 12 32 13 22	$\begin{array}{cccc} 02 & 24 \\ 03 & 38 \\ 04 & 50 \\ 05 & 58 \\ 06 & 55 \end{array}$
21 22 (b) 23 24 25	$ \begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{ccc} 06 & 15 \\ 06 & 59 \\ 07 & 36 \\ 08 & 08 \\ 08 & 37 \end{array}$	15 29 16 33 17 39 18 44 19 52	$\begin{array}{ccc} 06 & 34 \\ 07 & 15 \\ 07 & 49 \\ 08 & 18 \\ 08 & 44 \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 06 & 57 \\ 07 & 36 \\ 08 & 06 \\ 08 & 31 \\ 08 & 52 \end{array}$	14 35 15 50 17 05 18 21 19 36	07 28 08 01 08 26 08 46 09 02	14 25 15 38 16 56 18 15 19 32	07 39 08 12 08 36 08 53 09 06
26 27 28 29 30 C	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	09 03 09 28 09 55 10 22 10 55	$ \begin{array}{r} 20 & 58 \\ 22 & 05 \\ 23 & 15 \\ \hline 00 & 26 \end{array} $	$\begin{array}{cccc} 09 & 07 \\ 09 & 29 \\ 09 & 51 \\ 10 & 15 \\ 10 & 42 \end{array}$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 09 & 11 \\ 09 & 29 \\ 09 & 46 \\ 10 & 06 \\ 10 & 30 \end{array}$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 09 & 16 \\ 09 & 28 \\ 09 & 41 \\ 09 & 56 \\ 10 & 14 \end{array}$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	09 17 09 28 09 38 09 50 10 06
31	01 27	11 33	01 41	11 19	01 58	11 00	02 19	10 38	02 29	10 27
Feb. 1 2 3 4 5	02 40 03 52 04 57 05 54 06 40	12 21 13 19 14 25 15 39 16 53	02 58 04 11 05 19 06 12 06 55	$\begin{array}{cccc} 12 & 02 \\ 12 & 59 \\ 14 & 06 \\ 15 & 21 \\ 16 & 40 \end{array}$	03 18 04 35 05 41 06 34 07 12	11 41 12 35 13 42 15 01 16 24	03 45 05 06 06 12 07 00 07 33	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	03 57 05 21 06 27 07 13 07 43	10 59 11 49 12 58 14 23 15 55
6 () 7 8 9 10	07 19 07 50 08 18 08 43 09 08	18 05 19 14 20 19 21 22 22 23	07 29 07 57 08 19 08 43 09 05	17 56 19 09 20 18 21 25 22 29	$\begin{array}{cccc} 07 & 42 \\ 08 & 05 \\ 08 & 24 \\ 08 & 42 \\ 09 & 00 \end{array}$	$\begin{array}{cccc} 17 & 45 \\ 19 & 03 \\ 20 & 17 \\ 21 & 27 \\ 22 & 36 \end{array}$	07 57 08 14 08 29 08 41 08 54	$\begin{array}{cccc} 17 & 31 \\ 18 & 56 \\ 20 & 15 \\ 21 & 31 \\ 22 & 45 \end{array}$	08 03 08 19 08 30 08 40 08 50	17 26 18 52 20 14 21 33 22 50
11 12 13 14 15	09 35 10 02 10 33 11 07 11 52	$\begin{array}{c} 23 & 23 \\ \hline 00 & 24 \\ 01 & 23 \\ 02 & 22 \end{array}$	09 27 09 51 10 19 10 52 11 33	$\begin{array}{c} 23 & 33 \\ \hline 00 & 37 \\ 01 & 40 \\ 02 & 41 \end{array}$	09 17 09 38 10 02 10 33 11 10	$\begin{array}{cccc} 23 & 44 \\ \hline 00 & 52 \\ 02 & 00 \\ 03 & 04 \end{array}$	09 07 09 22 09 42 10 07 10 41	$\begin{array}{c} 23 & 58 \\ \hline 01 & 12 \\ 02 & 24 \\ 03 & 32 \end{array}$	09 02 09 15 09 31 09 54 10 27	$ \begin{array}{c} \hline 00 & 06 \\ 01 & 22 \\ 02 & 36 \\ 03 & 45 \end{array} $
16 17 18 19 20	$\begin{array}{cccc} 12 & 40 \\ 13 & 35 \\ 14 & 35 \\ 15 & 38 \\ 16 & 42 \end{array}$	$\begin{array}{ccc} 03 & 18 \\ 04 & 10 \\ 04 & 55 \\ 05 & 34 \\ 06 & 08 \end{array}$	12 20 13 17 14 19 15 24 16 32	03 38 04 29 05 13 05 50 06 21	$\begin{array}{ccccc} 11 & 57 \\ 12 & 54 \\ 13 & 58 \\ 15 & 08 \\ 16 & 19 \end{array}$	$\begin{array}{cccc} 04 & 02 \\ 04 & 52 \\ 05 & 33 \\ 06 & 07 \\ 06 & 34 \end{array}$	11 26 12 24 13 32 14 46 16 04	$\begin{array}{cccc} 04 & 32 \\ 05 & 22 \\ 06 & 00 \\ 06 & 29 \\ 06 & 51 \end{array}$	11 11 12 09 13 19 14 37 15 57	04 47 05 36 06 14 06 39 06 59
21 @ 22 23 24 25	$\begin{array}{cccc} 17 & 46 \\ 18 & 50 \\ 19 & 55 \\ 21 & 00 \\ 22 & 09 \end{array}$	06 39 07 06 07 32 07 59 08 26	$\begin{array}{cccc} 17 & 39 \\ 18 & 47 \\ 19 & 55 \\ 21 & 05 \\ 22 & 17 \end{array}$	06 47 07 11 07 33 07 56 08 19	$\begin{array}{cccc} 17 & 31 \\ 18 & 44 \\ 19 & 56 \\ 21 & 11 \\ 22 & 27 \end{array}$	06 56 07 16 07 35 07 52 08 12	17 21 18 39 19 58 21 18 22 41	07 08 07 23 07 36 07 49 08 02	17 17 18 38 19 58 21 21 22 46	07 14 07 25 07 36 07 47 07 59
26 27 28 C	$\frac{23}{00} \frac{19}{31}$	08 57 09 33 10 17	$\frac{23}{00} \frac{32}{48}$	08 46 09 20 09 59	$\frac{23}{01} \frac{47}{07}$	08 34 09 02 09 39	$\begin{array}{c} \hline 00 & 06 \\ 01 & 33 \end{array}$	08 19 08 42 09 13	$\begin{array}{c} \hline 00 & 15 \\ 01 & 45 \end{array}$	08 13 08 32 09 00

Date	Latitude 35° Moon Rise Set	Latitude 40° Moon Rise Set	Latitude 45° Moon Rise Set	Latitude 50° Moon Rise Set	Latitude 52° Moon Rise Set
Mar. 1 2 3 4 5	h m h m 01 42 11 10 02 49 12 12 03 47 13 22 04 36 14 36 05 15 15 46	h m h m 02 01 10 51 03 08 11 53 04 05 13 04 04 51 14 19 05 27 15 34	h m h m 02 25 10 27 03 33 11 29 04 28 12 42 05 10 14 02 05 41 15 21	h m h m 02 54 09 57 04 04 10 57 04 56 12 14 05 34 13 39 05 59 15 06	h m h m 03 09 09 42 04 19 10 42 05 11 12 00 05 47 13 28 06 08 14 58
6 7 0 8 9 10	0548165506171800064319040708200707342108	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	06 25 16 25 06 36 17 48 06 47 19 08 06 57 20 26 07 08 21 43
11 12 13 14 15	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
16 17 18 19 20	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
21 22 23 @ 24 25	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
26 27 28 29 30 C	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
31	02 33 12 24	02 50 12 08	03 11 11 48	03 36 11 23	03 48 11 10
Apr. 1 2 3 4 5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	03 45 13 07 04 11 14 24 04 32 15 38 04 51 16 50 05 08 17 59	04 04 12 50 04 25 14 13 04 41 15 32 04 55 16 49 05 07 18 04	04 14 12 41 04 31 14 06 04 45 15 29 04 56 16 48 05 06 18 06
6 (b) 7 8 9 10	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
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21 (D) 22 23 24 25	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
26 27 28 C 29 30	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$

Date	Latitude 35° Moon Rise Set	Latitude 40° Moon Rise Set	Latitude 45° Moon Rise Set	Latitude 50° Moon Rise Set	Latitude 52° Moon Rise Set
May 1 2 3 4 5	h m h m 02 49 14 44 03 15 15 45 03 39 16 46 04 05 17 46 04 32 18 47	h m h m 02 53 14 42 03 15 15 46 03 36 16 50 03 58 17 55 04 22 18 59	h m h m 02 57 14 40 03 15 15 49 03 32 16 57 03 50 18 05 04 10 19 14	h m h m 03 02 14 37 03 15 15 51 03 27 17 05 03 40 18 18 03 56 19 33	h m h m 03 04 14 36 03 14 15 53 03 24 17 09 03 36 18 25 03 48 19 41
6 7 8 9 10	05 03 19 48 05 39 20 47 06 21 21 44 07 08 22 35 08 01 23 21	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
11 12 13 14 15	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c} 08 \ 41 \ \ \ 09 \ 44 \ 00 \ 16 \ 10 \ 48 \ 00 \ 47 \ 11 \ 53 \ 01 \ 13 \ 12 \ 58 \ 01 \ 37 \ \end{array} $	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
16 17 18 19 20	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
21 (D) 22 23 24 25	20 06 04 44 21 19 05 40 22 21 06 46 23 11 07 59 23 51 09 13	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
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THE PLANETS FOR 1951

By C. A. Chant

THE SUN

During the present sun-spot cycle there has been remarkable activity on the sun. The maximum occurred about March 26, 1947, and the activity is still notable. In 1950 the sun was seldom without spots on its surface and they were often accompanied by striking auroral displays. Very probably this condition will continue throughout 1951.

MERCURY

Mercury is exceptional in many ways. It is the planet nearest the sun and travels fastest in its orbit, its speed varying from 23 mi. per sec. at aphelion to 35 mi. per sec. at perihelion. The amount of heat and light from the sun received by it per square mile is, on the average, 6.7 times the amount received by the earth. Its period of rotation on its axis is believed to be the same as its period of revolution about the sun, which is 88 days.

Mercury's orbit is well within that of the earth, and the planet, as seen from the earth, appears to move quickly from one side of the sun to the other several times in the year. Its quick motion earned for it the name it bears. Its greatest elongation (i.e., its maximum angular distance from the sun) varies between 18° and 28°, and on such occasions it is visible to the naked eye for about two weeks.

When the elongation of Mercury is east of the sun it is an evening star, setting soon after the sun. When the elongation is west, it is a morning star and rises shortly before the sun. Its brightness when it is treated as a star is considerable but it is always viewed in the twilight sky and one must look sharply to see it.

The most suitable times to observe Mercury are at an eastern elongation in the spring and at a western elongation in the autumn. The dates of greatest elongation this year, together with the planet's separation from the sun and its stellar magnitude, are given in the following table:

Elong. East—Evening Star			Elong. West—Morning Star			
Date	Distance	Mag.	Date	Distance	Mag.	
Apr. 5	19°	-0.2	Jan. 23	25°	+0.1	
Aug. 3	27°	+0.6	May 22	2 5°	+0.8	
Nov. 28	22°	+0.5	Sept. 16	18°	-0.1	

Maximum Elongations of Mercury during 1951

The most favourable elongations to observe are: in the evening, April 5; in the morning, Sept. 16. At these times Mercury is about 80 million miles from the earth and in a telescope looks like a half-moon about 7'' in diameter.

VENUS

Venus is the next planet in order from the sun. In size and mass it is almost a twin of the earth. Venus being within the earth's orbit, its apparent motion is similar to Mercury's but much slower and more stately. The orbit of Venus is almost circular with radius of 67 million miles, and its orbital speed is 22 miles per sec.

On Jan. 1, 1951, Venus is an evening star. It crosses the meridian about an hour after the sun, but as its declination is almost 23° S. it is not well placed for observers in the northern hemisphere. It is separating from the sun and on June 25 reaches greatest elong. E., 45° 25'. Then it moves in towards the sun. It attains greatest brilliancy, stellar mag. -4.2, on July 29 and reaches inferior conjunction with the sun on Sept. 3. The planet now becomes a morning star, rapidly separating from the sun. On Oct. 10 it attains greatest brilliancy, stellar mag. -4.3; and reaches greatest elong. W., 46° 39', on Nov. 14. It will be a morning star the rest of the year.

In the month of February the planets Venus, Mars and Jupiter have close conjunctions with each other and with the moon. For details consult page 32, the Sky for February, and the map of the path of Mars.

With the exception of the sun and moon, Venus is the brightest object in the sky. Its brilliance is largely due to the dense clouds which cover the surface of the planet. They reflect well the sun's light; but they also prevent the astronomer from detecting any solid object on the surface of the body. If such could be observed it would enable him to determine the planet's rotation period. It is probably around 30 days.

MARS

The orbit of Mars is outside that of the earth and consequently its planetary phenomena are quite different from those of the two inferior planets discussed above. Its mean distance from the sun is 141 million miles and the eccentricity of its orbit is 0.093, and a simple computation shows that its distance from the sun ranges between 128 and 154 million miles. Its distance from the earth varies from 35 to 235 million miles and its brightness changes accordingly. When Mars is nearest it is conspicuous in its fiery red, but when farthest away it is no brighter than Polaris. Unlike Venus, its atmosphere is very thin, and features on the solid surface are distinctly visible. Utilizing them its rotation period of 24h. 37m. has been accurately determined.

The sidereal, or true mechanical, period of revolution of Mars is 687 days; and the synodic period (for example, the interval from one opposition to the next one) is 780 days. This is the average value; it may vary from 764 to 810 days. The planet was in opposition on Mar. 23, 1950; the next opposition is on May 1, 1952; and the next close opposition is on Sept. 10, 1956. For its position among the stars see the map.

JUPITER

Jupiter is the giant of the family of the sun. Its mean diameter is 87,000 miles and its mass is $2\frac{1}{2}$ times that of all the rest of the planets combined! Its mean distance is 483 million miles and the revolution period is 11.9 years. This





planet is known to possess 11 satellites, two of them discovered in 1938 (see p. 59). Not so long ago it was generally believed that the planet was still cooling down from its original high temperature, but from actual measurements of the radiation from it to the earth it has been deduced that the surface is at about -200° F. The spectroscope shows that its atmosphere is largely ammonia and methane.

Jupiter is a fine object for the telescope. Many details of the surface as well as the flattening of the planet, due to its short rotation period, are visible, and the phenomena of its satellites provide a continual interest.

On Jan. 1, 1951, Jupiter crosses the meridian at 3.45 p.m. and is an evening star in the constellation Aquarius (see map). The sun moves over to the planet and they are in conjunction on March 11, and Jupiter becomes a morning star. It then separates from the sun until Oct. 2 when it comes to opposition and is on the meridian at midnight. At this time its distance from the earth is 366,800,000 mi. (see p. 45) and its stellar magnitude -2.5. On Dec. 31 it crosses the meridian at about 5.45 p.m.

SATURN

Saturn was the outermost planet known until modern times. In size it is a good second to Jupiter. In addition to its family of nine satellites, this planet has a unique system of rings, and it is one of the finest of celestial objects in a good telescope. The plane of the rings makes an angle of 27° with the plane of the planet's orbit, and twice during the planet's revolution period of $29\frac{1}{2}$ years the rings appear to open out widest; then they slowly close in until, midway between the maxima, the rings are presented edgewise to the sun or the earth, at which times they are invisible. The rings were edgewise in 1937 and 1950, and at maximum in 1944. For the next few years they will be gradually opening out.

The planet is in the constellation Virgo (see map). On March 20 it is in opposition to the sun and is visible all night. Its stellar magnitude then is +0.7, slightly less bright than Rigel. On June 17 it is in quadrature with the sun and is on the meridian at sunset. On Sept. 29 it is in conjunction with the sun. On Dec. 19, it is in close conjunction with Mars (see p. 53).



URANUS

Uranus was discovered in 1781 by Sir William Herschel by means of a 6¼-in. mirror-telescope made by himself. The object did not look just like a star and he observed it again four days later. It had moved amongst the stars, and he assumed it to be a comet. He could not believe that it was a new planet. How-



ever, computation later showed that it was a planet nearly twice as far from the sun as Saturn. Its period of revolution is 84 years and it rotates on its axis in about 11 hours. Its five satellites are visible only in a large telescope. The fifth satellite was discovered by G. P. Kuiper in 1948 at the McDonald Observatory (see p. 59).

As shown by the chart, Uranus in 1951 is still in Gemini. It was in opposition to the sun on Dec. 29, 1950, but as its synodic period is 369.66 days there will be no opposition in 1951; it will occur on Jan. 3, 1952.

NEPTUNE

Neptune was discovered in 1846 after its existence in the sky had been predicted from independent calculations by Leverrier in France and Adams in England. This discovery was a crowning demonstration of the correctness of Newton's law of gravitation. It caused a sensation at the time. The planet's



distance from the sun is 2800 million miles and its period of revolution is 165 years. A satellite was discovered in 1846, soon after the planet. A second satellite was discovered by G. P. Kuiper at the McDonald Observatory on May 1, 1949. Its magnitude is about 19.5, its period about 2 years, and diameter about 200 miles. It is named Nereid.

During 1951 Neptune is still in the constellation Virgo. It is in opposition to the sun on April 8. Its stellar magnitude is +7.7 and hence it is too faint for the naked eye. In the telescope it shows a greenish tint and a diameter of 2''.5. It is in conjunction with the sun on Oct. 13.

PLUTO

Pluto, the most distant known planet, was discovered at the Lowell Observatory in 1930. Its mean distance from the sun is 3666 million miles and its revolution period is 248 years. It appears as a 15th mag. star in the constellation Cancer. It is in opposition to the sun on Feb. 8, 1951, at which time its astrometric position is R.A. 9^{h} 36^{m} . 0, Dec. $+23^{\circ}$ 32'.

ECLIPSES, 1951

In 1951 there will be the least possible number of eclipses, *two*, both of the sun. Both of these eclipses are annular.

I. An Annular Eclipse of the Sun, March 7, 1951. Visible in Canada and the United States only as a slight partial eclipse at sunset from southern Ontario, and from the southern and eastern states. The central path starts between Australia and New Zealand, crosses the South Pacific, Pitcairn Island, Nicaragua and Costa Rica, and ends in the Caribbean Sea at sunset. The maximum duration of the annular phase will be 1 min. 38 sec. The G.C.T. of conjunction in right ascension will be March 7th, 20h 38m 37.5s.

II. An Annular Eclipse of the Sun, September 1, 1951, visible as a partial eclipse at sunrise from eastern Canada and the United States, and briefly as an annular eclipse of the rising sun from eastern Virginia. The central path starts at sunrise in Virginia, crosses the North Atlantic Ocean, French West Africa, The Gold Coast, Gulf of Guinea, the Congo, Northern Rhodesia, Mozambique, and ends in Madagascar at sunset. The duration along the central line of the path is remarkably constant, ranging from 2 min. 33 sec. to 2 min. 43 sec. The G.C.T. of conjunction in right ascension will be September 1st, 12h 42m 1.6s.

THE SKY MONTH BY MONTH

By J. F. HEARD

THE SKY FOR JANUARY, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45°N.

The Sun—During January the sun's R.A. increases from 18h 42m to 20h 55m and its Decl. changes from 23° 05' S. to 17° 23' S. The equation of time changes from -3m 06s to -13m 31s. The earth is in perihelion or nearest the sun on the 1st. For changes in the length of the day, see p. 11.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 18.

Mercury on the 15th is in R.A. 18h 10m, Decl. 20° 46' S. and transits at 10.34. After inferior conjunction on the 1st it is a morning star, coming to greatest western elongation on the 23rd at which time it may be seen low in the south-east just before sunrise.

Venus on the 15th is in R.A. 20h 47m, Decl. 19° 26' S. and transits at 13.13. It is an evening star very low in the south-west at sunset. In a telescope it appears nearly fully illuminated.

Mars on the 15th is in R.A. 21h 47m, Decl. 14° 31' S. and transits at 14.12. It is in Capricornus and Aquarius and may be seen low in the south-west after sunset:

Jupiter on the 15th is in R.A. 22h 38m, Decl. 9° 45' S. and transits at 15.01. It is prominent in the south-west for a few hours after sunset.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 12h 12m, Decl. 1° 12' N. and transits at 4.36. It is in Virgo west of Spica, rising about midnight. On the 13th it is stationary in R.A. and begins to retrograde, or move westward among the stars. The rings are "thin" this year, their plane being within a few degrees of the line of sight.

Uranus on the 15th is in R.A. 6h 30m, Decl. 23° 36' N. and transits at 22.50

Neptune on the 15th is in R.A. 13h 15m, Decl. 6° 08' S. and transits at 5.38.

Pluto-For information in regard to this planet, see p. 29.

<u></u>				JANUARY	Min.	Config. of Juniter's
				75th Meridian Civil Time	of Algol	Sat. 19h 00m
	d	h	m		h m	ı
Mon.	1	0	11	C Last Quarter		40213
		15		σ ['] ₿⊙ Inferior		
		19	41			
		23		\oplus in Perihelion. Dist. from \odot , 91, 345,000 mi.		
Tue.	2	22		Q in Aphelion		14023
Wed.	3			Quadrantid meteors	15 22	2 23014
Thur.	4					32104
Fri.	5					30124
Sat.	6	8		Moon in Perigee. Dist. from \oplus , 223,500 mi	12 12	2 3024*
		19	08	σ₿Œ₿8°12′N		
Sun.	7	15	10	New Moon		21034
Mon.	8	2		§ Greatest Hel. Lat. N.,	1	0134*
		13	27	ϭʹ♀ € ♀ 2° 38′ N	1	
Tue.	9	21	14	of o ⁷ € o ⁷ 1° 36′ N	09 01	1 10234
Wed.	10	8		Ψ Ο		23041
Thur.	11	2	31	σ 24 0° 13′ N		32140
Fri.	12	10		§ Stationary in R.A.	05 50	34012
Sat.	13	0		b Stationary in R.A.		43102
Sun.	14	19	23	First Ouarter.		d42O3
Mon.	15			~	02 39	9 4013*
Tue.	16					41023
Wed.	17				23 29	9 42031
Thur.	18	9		Moon in Apogee. Dist. from ⊕. 252.000 mi.		34210
Fri.	19			F-8		3012*
Sat.	20	20	38	්ර බ ⊈ බ 4° 34′ S	20 18	3 31024
Sun.	21	16		Ψ Stationary in R.A		20134
Mon.	22	23	47	Full Moon		2034*
Tue.	23	18		g Greatest elongation W., 24° 31'	17 0	7 10234
Wed.	24					20314
Thur.	25	11		Q Greatest Hel. Lat. S.		32104
Fri.	26				13 5	7 30214
Sat.	27	18	29	♂ b € b 3° 58′ N		31042
Sun.	28				1	2401*
Mon.	29	1	49	σΨ € Ψ 4° 43′ N	10 4	6 42103
Tue.	30	10	13	Last Quarter		41023
Wed.	31	12		ይ in የን		40213

ASTRONOMICAL PHENOMENA MONTH BY MONTH

By Ruth J. Northcott

Explanation of symbols and abbreviations on p. 4, of time on p. 8.

THE SKY FOR FEBRUARY, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During February the sun's R.A. increases from 20h 55m to 22h 45m and its Decl. changes from $17^{\circ} 23'$ S. to $7^{\circ} 59'$ S. The equation of time changes from -13m 31s to a maximum of -14m 20s on the 12th and then to -12m 39s at the end of the month. For changes in the length of the day, see p. 11.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 18.

Mercury on the 15th is in R.A. 20h 43m, Decl. $19^{\circ} 52'$ S. and transits at 11.08. It is a morning star but poorly placed for observation.

Venus on the 15th is in R.A. 23h 17m, Decl. 6° 08' S. and transits at 13.40. It is an evening star seen low in the west just after sunset. There is an occultation on the 7th, in the evening in eastern Canada, in the afternoon in western Canada. Conjunction with Jupiter is on the 11th and with Mars on the 15th. See Mars.

Mars on the 15th is in R.A. 23h 19m, Decl. $5^{\circ}20'$ S. and transits at 13.41. It is in Aquarius and Pisces and may be seen for a short time after sunset low in the south-west. Conjunction with Jupiter is on the 7th and with Venus on the 15th. These are close conjunctions and the closeness and relative positions of these three planets at this time will be most interesting.

Jupiter on the 15th is in R.A. 23h 04m, Decl. 7° 05' S. and transits at 13.25. It is very low in the south-west at sunset, setting soon after. See also Venus and Mars.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 12h 09m, Decl. 1° 43' N. and transits at 2.31. It is in Virgo west of Spica, rising in the late evening and visible the rest of the night.

Uranus on the 15th is in R.A. 6h 25m, Decl. 23° 39' N. and transits at 20.44.

Neptune on the 15th is in R.A. 13h 14m, Decl. 6° 03' S. and transits at 3.36.

Pluto—For information in regard to this planet, see p. 29.

FEBRUARY 75th Meridian Civil Time						Config. of Jupiter's Sat. 18h 45m
	d	h	m		h m	n
Thur.	1				07 35	6 42310
Fri.	2					43021
Sat.	3	10		Moon in Perigee. Dist. from \oplus , 226,800 mi		43102
Sun.	4	10	09	୪ ଓ ଏ ପ୍ର° 53′ N	04 2 4	l 4201*
Mon.	5					2103*
Tue.	6	2	54	New Moon		dO243
Wed.	7	14		ഗ്രീ 24 ഗീ 0° 10′ N	01 14	l 01234
		17	11	ଟ ହ ଏ ହେ 0° 36′ S		
Thur.	8	0	07	σ 24 0° 30′ S		23104
		0	34	ර ් ℃ ් 0° 21′S		
		1		சி சி ⊙		
Fri.	9				22 03	3 3014*
Sat.	· 10	19		ξ in Aphelion		31024
Sun.	11	10		σ ♀ 24 ♀ 0° 26′ S		23014
Mon.	12				18 52	2 21034
Tue.	13	15	55	First Quarter		01423
Wed.	14			~		4023*
Thur.	15	5		Moon in Apogee. Dist. from \oplus , 251,400 mi.	15 42	2 42130
		23		ດ′♀ ດ ⁷ ♀ 0° 35′ S		
Fri.	16					43201
Sat.	17	2	48	ර ඊ € 6 4° 42′ S		43102
Sun.	18				12 3	1 43201
Mon.	19					42103
Tue.	20					40123
Wed.	21	16	12	Full Moon	09 20	0 41023
Thur.	22					
Fri.	23	22	19	♂ þ @ þ 4° 03′ N		
Sat.	24				06 10	0
Sun.	25	6	58	σΨ@ Ψ4°50′ N		-
Mon.	26					
Tue.	27				02 59	9
Wed.	28	17	59	Last Quarter		

Explanation of symbols and abbreviations on p. 4, of time on p. 8.

Jupiter being near the sun, phenomena of the satellites are not given from February 22nd to May 1st.

THE SKY FOR MARCH, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun-During March the sun's R.A. increases from 22h 45m to 0h 38m and its Decl. changes from 7° 59' S. to 4° 09' N. The equation of time changes from -12m 39s to -4m 16s. On the 21st at 5.26 E.S.T. the sun crosses the equator on its way north, enters the sign of Aries and spring commences. This is the vernal equinox. There is an annular eclipse of the sun on the 7th. For changes in the length of the day, see p. 12.

The Moon-For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 19.

Mercury on the 15th is in R.A. 23h 51m, Decl. 2° 16' S. and transits at 12.26. It is in superior conjunction on the 11th and is too close to the sun for observation until the end of the month when it becomes a good evening star.

Venus on the 15th is in R.A. 1h 23m, Decl. 8° 17' N. and transits at 13.56. It is an evening star seen low in the west after sunset.

Mars on the 15th is in R.A. 0h 38m, Decl. 3° 31' N. and transits at 13.10. It is in Pisces, very low in the west at sunset.

Jupiter on the 15th is in R.A. 23h 29m, Decl. 4° 28' S. and transits at 12.00. It is too near the sun for observation. Conjunction is on the 11th.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 12h 02m, Decl. 2° 33' N. and transits at 0.34. It is in Virgo west of Spica rising in the early evening and visible all night. Opposition is on the 20th.

Uranus on the 15th is in R.A. 6h 24m, Decl. 23° 40' N. and transits at 18.53.

Neptune on the 15th is in R.A. 13h 12m, Decl. 5° 49' S. and transits at 1.44

Pluto-For information in regard to this planet, see p. 29.
MARCH Mig	n.
75th Meridian Civil Time Alg	ol
d h m hn	n
Thur. 1 23 -	48
Fri. $2 2 $ Moon in Perigee. Dist. from \oplus , 229,800 mi.	
Sat. $3 \begin{vmatrix} 3 \\ 3 \end{vmatrix} = 0$ Greatest Hel. Lat. S	
Sun. 4 20	38
Mon. 5	
Tue. 6	
Wed. 7 10 22 $\sigma \notin \mathbb{Q} \oplus 1^{\circ} 30' \text{ S}.$	54. A
Annular eclipse of \bigcirc . See p. 29	27
15 50 W New Moon	
22 02 32 02 02 02 02 02 02 02 02 02 02 02 02 02	
Thur. 8	
Fri. 9 4 36 $\sigma \sigma^2 \mathbb{C}$ $\sigma^2 2^{\circ} 23' S$	
Sat. 10 1 01 $\sigma' \notin \mathbb{C}$ φ 3° 24' S 14	16
Sun. 11 5 0 \$ 24 \$ 0° 37' S	
$5 \circ \circ \circ \circ \circ$ Superior	
$ 12 $ $\sigma' 24 \odot \dots$	
Mon. 12	
Tue. 13	06
Wed. 14 6 8 Stationary in R.A. 8	
Thur. 15 1 Moon in Apogee. Dist. from \oplus , 251,200 mi.	
12 40 D First Quarter	
Fri. 16 10 35 5 🕅 👶 4° 45' S 07	55
Sat. 17	
Sun. 18	
Mon. 19 04 4	44
Tue. 20 5 $0^{\circ} b^{\circ}$ Dist. from \oplus , 788,600,000 mi.	
Wed. 21 5 26 \odot enters γ . Spring commences. Long. of $\odot, 0^{\circ}$	
Thur. 22 3 β in ω 01 :	33
18 Q in Q	
Fri. 23 2 54 54 b C b 3° 53' N	-
5 50 🕲 Full Moon	
Sat. 24 3 3 $\forall \Psi \oplus \Psi $	22
Sun. 25	
Mon. 26 4 $\sigma \notin \sigma^{-1}$	
19 § in Perihelion	
Tue. 27 4 Moon in Perigee. Dist. from \oplus , 228,600 mi. 19	12
Wed. 28	
Thur. 29	
Fri. 30 0 35 C Last Quarter 16	01

Explanation of symbols and abbreviations on p. 4, of time on p. 8. Jupiter being near the sun, phenomena of the satellites are not given from February 22 to May 1.

THE SKY FOR APRIL, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During April the sun's R.A. increases from 0h 38m to 2h 30m and its Decl. changes from 4° 09' N. to 14° 46'N. The equation of time changes from -4m 16s to +2m 48s, being zero on the 15th; that is, the apparent sun changes from being east of the mean sun to being west of it. For changes in the length of the day, see p. 12.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 19.

Mercury on the 15th is in R.A. 2h 21m, Decl. 17° 14' N. and transits at 12.48. It is at greatest eastern elongation on the 5th; around about this time it is a splendid evening star, being 18° above the western horizon at sunset. However, it rapidly approaches the sun and by the 24th it is in inferior conjunction.

Venus on the 15th is in R.A. 3h 48m, Decl. 21° 14' N. and transits at 14.20. It is a good evening star prominent in the western sky in the early evening.

Mars on the 15th is in R.A. 2h 05m, Decl. $12^{\circ} 29'$ N. and transits at 12.35. It is too close to the sun for observation.

Jupiter on the 15th is in R.A. 23h 56m, Decl. 1° 34' S. and transits at 10.25. It is too near the sun for easy observation.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 11h 53m, Decl. 3° 28' N. and transits at 22.19. It is in Virgo west of Spica, well up in the east at sunset and visible all night.

Uranus on the 15th is in R.A. 6h 26m, Decl. 23° 38' N. and transits at 16.53.

Neptune on the 15th is in R.A. 13h 09m, Decl. 5° 30' S. and transits at 23.35.

				APRIL	Min.				
75th Meridian Civil Time									
	d	h	m		hm	1			
Sun.	1	Į		•••••••••••••••••••••••••••••••••••••••		1			
Mon.	2				12 50				
Tue.	3								
Wed.	4	18	35	σ 24 € 24 1° 56′ S					
Thur.	5	15		§ Greatest elongation E., 19° 12'	09 39				
Fri.	6	1		§ Greatest Hel. Lat. N.					
		5	52	New Moon					
Sat.	7	7	25	o′o¹ € o¹ 3° 54′ S					
		21	26	σ Ϩ 🕼 🛛 Ϩ 1° 01′ S					
Sun.	8	15	1	$\circ^{\circ} \Psi \odot$ Dist. from \oplus . 2.722.000.000 mi	06 28				
Mon.	9	7	53	$\sigma \neq 0$ $\varphi = 3^{\circ} 57' S$	00 -0				
Tue.	10								
Wed.	11	20		Moon in Apogee. Dist. from (#) 251 600 mi	03 18				
Thur.	12	19	22	♂ å Ø å 4° 39′ S	00 10				
Fri.	13								
Sat.	14	7	55	First Quarter	00.07				
Sati		17		8 Stationary in R A	00 01	1.1			
Sun	15					4			
Mon	16				20 56				
Tue	17			•••••••••••••••••••••••••••••••••••••••	20 30				
Wed	18								
Thur	10	3		α 8 α ⁷ 8 2° 51′ N	17 45				
r nur.	10	q	13	$\alpha = 0$ $\psi = 2^{-51}$ N	17 40				
Fri	20	21	10	\sim tu σ tu $a^{\circ} a^{2}$ N					
Sat	20	16	30						
Sat.	21	10	00	L vrid meteors					
Sun	9 9				14 94				
Mon	22	18		Moon in Parizon Dist from (225 200 mi	14 04				
Tuo	20 94	10		$\sim 8 \odot$ Inferior					
Wod	41 95	20		$0 \neq 0$ in Dovibation	11 09				
Thum	20 96	9			11 23				
Thur.	20 97			•••••••••••••••••••••••••••••••••••••••					
rri. Sat	21	7	17	A Lost Ouenten	00 10				
Sat.	28		11	Last Quarter	08 12				
Sun.	29	11		۷ in ۵۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰۰					
Mon.	30			·········					

Explanation of symbols and abbreviations on p. 4, of time on p. 8. Jupiter being near the sun, phenomena of the satellites are not given from February 22 to May 1.

THE SKY FOR MAY, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During May the sun's R.A. increases from 2h 30m to 4h 32m and its Decl. changes from 14° 46′ N. to 21° 55′ N. The equation of time changes from +2m 48s to a maximum of +3m 45s on the 15th and then to +2m 37s at the end of the month. For changes in the length of the day, see p. 13.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 20.

Mercury on the 15th is in R.A. 1h 56m, Decl. 8° 31' N. and transits at 10.27. It is a morning star, reaching greatest western elongation on the 22nd, but this is an unfavourable elongation.

Venus on the 15th is in R.A. 6h 20m, Decl. 25° 43' N. and transits at 14.53. It is a prominent evening star dominating the western sky.

Mars on the 15th is in R.A. 3h 32m, Decl. 19° 13' N. and transits at 12.03. It is too close to the sun for observation. Conjunction is on the 22nd.

Jupiter on the 15th is in R.A. 0h 20m, Decl. 0° 56' N. and transits at 8.51. It is a prominent morning star low in the east before sunrise.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 11h 48m, Decl. 3° 57' N. and transits at 20.16. It is in Virgo about half-way between Spica and Regulus. It is well up in the southern sky at sunset and sets after midnight. On the 29th it is stationary in R.A. and resumes direct, or eastward, motion.

Uranus on the 15th is in R.A. 6h 30m, Decl. 23° 35' N. and transits at 15.00.

Neptune on the 15th is in R.A. 13h 06m, Decl. 5° 14' S. and transits at 21.34.

				MAY	Min. of	Config of Jupiter's
				75th Meridian Civil Time	Algol	Sat. 4h 15m
	d	h	m		h m	
Tue.	1	10		σ^{\uparrow} in Ω	05 01	32104
Wed.	2	13	03	♂ 24 ℃		20314
Thur.	3			•••••••••••••••••••••••••••••••••••••••		10234
Fri.	4			Eta Aquarid Meteors	01 50	02134
-		17	07	σ ⊈ € 5° 13′ S		
Sat.	5	20	35	New Moon		21034
Sun.	6	7	52	໔ ♂ ℃ ♂ 4° 38′ S	22 39	3014*
Mon.	7	6		ØStationary in R.A		3024*
Tue.	8					d32O4
Wed.	9	12		Moon in Apogee. Dist. from \oplus , 252,200 mi	19 28	23041
		12	08	σ ♀ ℂ ♀ 2° 44′ S		
		18		۵ in Aphelion		
Thur.	10	4	27	ା ଦ ବି ଐ ି ଶି 4° 27′ S		41023
Fri.	11			• • • • • • • • • • • • • • • • • • • •		40213
Sat.	12			•••••••••••••••••••••••••••••••••••••••	16 17	42103
Sun.	13					4301*
Mon.	14	0	32	First Quarter		43102
Tue.	15				13 06	d4320
Wed.	16	16	58	σ þ 🕼 þ 3° 41′ N		42301
Thur.	17	0		$\sigma \neq \delta \qquad \varphi 2^{\circ} 05' \text{ N}$		41023
		6		Q Greatest Hel. Lat. N		1.4.6
Fri.	18	7	00	σΨ 🕼 Ψ 4° 45′ N	09 55	04123
Sat.	19			•••••••••••••••••••••••••••••••••••••••		21034
Sun.	20			·····		23014
Mon.	21	0	45	Full Moon	06 44	31024
-		23		Moon in Perigee. Dist. from \oplus , 222,700 mi		
Tue.	22	9		ơ ơ' ⊙		d3O14
	~~	12		φ Greatest elongation W., 25° 24'		
Wed.	23			•••••••••••••••••••••••••••••••••••••••		2304*
I hur.	24			•••••••••••••••••••••••••••••••••••••••	03 33	10234
Fri.	25			•••••••••••••••••••••••••••••••••••••••		01243
Sat.	26			·····		21043
Sun.	27	15	17	Last Quarter	00 22	d42O1
Mon.	28	-		L Cont t D A		43102
1 ue.	29	20		P Stationary in R.A.	21 11	43021
Wed.	30	2				4230*
T 1		ъ	22	σ 4 @ 4 3° 30′ S		
1 hur.	31					d4O23

THE SKY FOR JUNE, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During June the sun's R.A. increases from 4h 32m to 6h 36m and its Decl. changes from $21^{\circ} 55'$ N. to $23^{\circ} 27'$ N. at the solstice on the 22nd and then to $23^{\circ} 11'$ N. at the end of the month. The equation of time changes from +2m29s to -3m 28s, being zero on the 14th; that is, the apparent sun changes from being west of the mean sun to being east of it. For changes in the length of the day, see p. 13.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 20.

Mercury on the 15th is in R.A. 4h 37m, Decl. 21° 26' N. and transits at 11.10. It is poorly placed for observation all month. Superior conjunction is on the 25th.

Venus on the 15th is in R.A. 8h 44m, Decl. 20° 19' N. and transits at 15.15. It is a good evening star dominating the western sky. Greatest eastern elongation is on the 25th. At this time it appears like a half-moon in the telescope.

Mars on the 15th is in R.A. 5h 03m, Decl. 23° 16' N. and transits at 11.33. It is too close to the sun for observation.

Jupiter on the 15th is in R.A. 0h 40m, Decl. 2° 58' N. and transits at 7.08. It rises shortly after midnight and dominates the eastern sky the rest of the night.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 11h 48m, Decl. 3° 50' N. and transits at 18.15. It is in Virgo about half-way between Spica and Regulus. It is well past the meridian at sunset and sets before midnight.

Uranus on the 15th is in R.A. 6h 38m, Decl. 23° 29' N. and transits at 13.05.

Neptune on the 15th is in R.A. 13h 05m, Decl. 5° 04' S. and transits at 19.31.

			-	JUNE 75th Meridian Civil Time	Min. of Algol	Config. of Jupiter's Sat. 3h 30m	
	d	h	m		h m		
Fri.	1				18 00	40123	
Sat.	2	15	51	ර්⊈ී ਊ 7° 28′ S		42103	
Sun.	3					2031*	
Mon.	4	5	58	ර් ර් € ් ර් 4° 41′ S	14 48	31042	
		11	40	New Moon			
Tue.	5	20		Moon in Apogee. Dist. from \oplus , 252,600 mi.		30214	
Wed.	6	13	32	ර ී 🕻 👘 යි 4° 16′ S		32104	
Thur.	7				11 37	0314*	
Fri.	8	12	58	ଟ ହ ⊈ ଦୁ 1° 04′ S		0234*	
Sat.	9					21034	
Sun.	10				08 26	20134	
Mon.	11					31024	
Tue.	12	13	52	First Quarter		34021	
Wed.	13	1	30	σ þ 🕼 þ 3° 52′ Ν	05 15	34210	
Thur.	14	15	52	σΨ€ Ψ4° 56′ Ν		401**	
Fri.	15					4023*	
Sat.	16				02 04	d42O3	
Sun.	17	16		□ ♭ ⊙		42013	
Mon.	18	3		ម្ in Q	22 52	43102	
Tue.	19	7	36	Full Moon		34012	
		9		Moon in Perigee. Dist. from \oplus , 221,800 mi.			
		10		ర రై రి 0° 16′ S			
Wed.	20					32140	
Thur.	21				19 41	2014*	
Fri.	22	0	25	⊙enters⊗,Summer commences. Long.of⊙,90°		10234	
		18		§ in Perihelion			
Sat.	23					ddO34	
Sun.	24				16 30	20134	
Mon.	25	9		$\sigma \notin \odot$ Superior		13024	
		12		Q Greatest elongation E., 45° 25'			
Tue.	26	1	21	Last Quarter		30124	
		19	34	σ 24 € 24 4° 14′ S			
Wed.	27				13 18	32104	
Thur.	28	5		σ₿δ ₿1°10′N		23014	
Fri.	2 9	7		Ψ Stationary in R.A		14032	
Sat.	30				10 07	40213	
		•				•	

THE SKY FOR JULY, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During July the sun's R.A. increases from 6h 36m to 8h 41m and its Decl. changes from 23° 11' N. to 18° 17' N. The equation of time changes from -3m 28s to a maximum of -6m 24s on the 27th and then to -6m 17s at the end of the month. On the 4th the earth is in aphelion or farthest from the sun. For changes in the length of the day, see p. 14.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 21.

Mercury on the 15th is in R.A. 8h 57m, Decl. $+18^{\circ} 47'$ N. and transits at 13.31. It is poorly placed for observation all month.

Venus on the 15th is in R.A. 10h 29m, Decl. 9° 14' N. and transits at 14.59. It is an evening star seen low in the west. About the 5th it is interestingly close to Regulus. Greatest brilliancy is on the 29th when it reaches stellar magnitude -4.2.

Mars on the 15th is in R.A. 6h 32m, Decl. 23° 58' N. and transits at 11.03. It is too close to the sun for easy observation.

Jupiter on the 15th is in R.A. 0h 52m, Decl. 4° 05' N. and transits at 5.22. It rises before midnight and dominates the eastern and southern sky the rest of the night.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 11h 54m, Decl. 3° 07' N. and transits at 16.22. It is in Virgo west of Spica, well down in the south-west at sunset and setting a few hours later.

Uranus on the 15th is in R.A. 6h 46m, Decl. 23° 21' N. and transits at 11.15.

Neptune on the 15th is in R.A. 13h 05m, Decl. 5° 06' S. and transits at 17.33.

$\begin{array}{c c c c c c c c c c c c c c c c c c c $	03* 102 012
Sun.1	03* 102 012
Mon. 2 4 σ^{\prime} \circ	1O2 D12
23 Moon in Apogee. Dist. from \oplus , 252.600 mi.	D12
	D12
Tue. 3 0 § Greatest Hel. Lat. N 06 56 430	20
2 34 of o ⁷ C o ⁷ 4° 11′ S	20
22 39 of ô C ô 4° 10′ S	20
Wed. 4 2 48 W New Moon	
16 \oplus in Aphelion. Dist. from \odot , 94,459,000 mi.	
Thur. 5 2 40 $\sigma \notin \mathbb{C}$ ψ 1° 56' S	201
$ 10 \square 2 \odot \dots$	
Fri. 6 03 45 410	032
Sat. 7 40	123
Sun. 8 3 27 $\sigma \ \varphi \ (\ \varphi \ 0^{\circ} \ 34' \ S $ 210)43
Mon. 9 15 $\Box \Psi \odot$ 00 33 dO	34*
Tue. 10 10 36 $\sigma \not = 0$ $\phi = 00' N_{11} N_{12} N_{$	124
Wed. 11 23 27 $\sigma \Psi \oplus \Psi = 0^{\circ} 10^{\prime} \text{ N}_{\odot}$ 21 22 312	204
23 56 b First Quarter	
Thur. 12 7 9 in \mathfrak{N} 320	D14
Fri. 13	324
Sat. 14 18 10 01	234
Sun. 15 210	D43
Mon. 16 40	13*
Tue 17 18 Moon in Perigee. Dist. from \oplus . 222.600 mi. 14 59 430	02*
Wed. 18 14 17 @ Full Moon	120
Thur 19 18 $\sigma' \sigma' \delta = \sigma' 0^{\circ} 26' N_{\odot}$ 432	201
Fri 20 11 48 410	032
Sat. 21 40	123
Sun 22 42	103
Mon 23 08 36 420	013
Tue 24 7 19 σ 21 ℓ 21 4° 47' S	42*
Wed 25 13 59 \emptyset Last Quarter	104
Thur $26 \ 11 \ 8 \ in \ 9$	014
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	024
Sat 28 Delta Aquarid meteors	234
Sun 29 1 9 Greatest brilliancy	034
Mon $30 7 $ Moon in Apogee. Dist. from \oplus 252,200 mi 20	134
The 31 8 02 $\sigma \approx 0^{\circ}$ π° π°	024
22 39 σ	

THE SKY FOR AUGUST, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During August the sun's R.A. increases from 8h 41m to 10h 38m and its Decl. changes from $18^{\circ} 17'$ N. to $8^{\circ} 40'$ N. The equation of time changes from -6m 17s to -0m 20s. For changes in the length of the day, see p. 14.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 21.

Mercury on the 15th is in R.A. 10h 59m, Decl. 2° 40' N. and transits at 13.26. It is an evening star, being at greatest eastern elongation on the 3rd but the elongation is unfavourable. Inferior conjunction is on the 31st.

Venus on the 15th is in R.A. 11h 08m, Decl. 0° 43' S. and transits at 13.34. It is an evening star but very low in the west by sunset.

Mars on the 15th is in R.A. 8h 00m, Decl. $21^{\circ} 34'$ N. and transits at 10.29. It is in Gemini and Cancer. It rises about two hours before the sun and is not at all prominent.

Jupiter on the 15th is in R.A. 0h 54m, Decl. 4° 08' N. and transits at 3.22. Rising a few hours after sunset, it is very prominent the rest of the night. On the 4th it is stationary in R.A. and begins to retrograde, that is, move westward among the stars.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 12h 04m, Decl. 1° 56' N. and transits at 14.31. It is in Virgo west of Spica, well down in the west at sunset and setting soon after.

Uranus on the 15th is in R.A. 6h 53m, Decl. 23° 13' N. and transits at 9.20.

Neptune on the 15th is in R.A. 13h 06m, Decl. 5° 19' S. and transits at 15.33.

				AUGUST 75th Meridian Civil Time	Min. of Algol	Config. of Jupiter's Sat. 2h 00m
	d]	h	m		h m	
Wed.	1					d3O42
Thur.	2	17	3 9	New Moon		32401
Fri.	3	14		Greatest elongation E., 27° 19'	19 51	4102*
Sat.	4	16		24 Stationary in R.A		40123
Sun.	5	1	45	σ′ ἕ 🕻 🖞 2° 01′ S		41203
		15	54	ố ♀ ቒ ♀ 3° 41′ S		
		17		۵ in Aphelion		
Mon.	6	20	37	♂ 𝔥 𝔄	16 3 9	42013
Tue.	7					41302
Wed.	8	6	01	σΨ 🕻 Ψ 5° 21' Ν		43012
Thur.	9				13 28	3240*
Fri.	10	7	22	First Quarter		3104*
		19		Q Stationary in R.A		
Sat.	11					01324
Sun.	12			Perseid meteors.	10 16	12034
		20		♂ B⊙		
Mon.	13					20134
Tue.	14	23		Moon in Perigee. Dist. from \oplus , 225,000 mi.		d1024
Wed.	15	17		♀ in Aphelion	07 05	30124
Thur.	16	16		§ Stationary in R.A		3204*
		21	59	Full Moon		
Fri.	17					32104
Sat.	18				03 53	40132
Sun.	19					41203
Mon.	20	15	58	α 24 5° 01′ S		42013
Tue.	21				00 42	41032
Wed.	22					43012
Thur.	23				21 31	43210
Fri.	24	5	20	Last Quarter		d432O
Sat.	25					40132
Sun.	26	1		Greatest Hel. Lat. S	18 19	d14O3
		22		Moon in Apogee. Dist. from \oplus , 251,600 mi.		
Mon.	27	17	45	ି ପ ି ଏ ି ସଂ 08′ S		20143
Tue.	28					10234
Wed.	29	18	56	ା ଦ ଦି ⊈ ଦି 1° 52′ S	15 08	3 30124
Thur.	30					32104
Fri.	31	3		σ ♀ ⊙ Inferior		32014

Explanation of symbols and abbreviations on p. 4, of time on p. 8.

THE SKY FOR SEPTEMBER, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During September the sun's R.A. increases from 10h 38m to 12h 26m and its Decl. changes from 8° 40′ N. to 2° 47′ S. The equation of time changes from -0m 20s to +9m 57s, the apparent sun passing to the west of the mean sun on the 2nd. On the 23rd at 15.38 E.S.T. the sun crosses the equator moving southward, enters the sign of Libra, and autumn commences. There is an annular eclipse of the sun on the 1st. For changes in the length of the day, see p. 15.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 22. The full moon of September 15th is the Harvest Moon

Mercury on the 15th is in R.A. 10h 22m, Decl. 10° 21' N. and transits at 10.50-It is a morning star and near the time of greatest western elongation on the 16th it may be seen in the east near Regulus just before sunrise.

Venus on the 15th is in R.A. 10h 12m, Decl. $2^{\circ} 32'$ N. and transits at 10.37. Early in the month it is too close to the sun for observation, being in inferior conjunction on the 3rd. By the end of the month it may be seen as a morning star low in the east just before sunrise.

Mars on the 15th is in R.A. 9h 21m, Decl. 16° 43' N. and transits at 9.48. It is in Cancer and Leo, rising a few hours before the sun; not prominent.

Jupiter on the 15th is in R.A. 0h 45m, Decl. 3° 03' N. and transits at 1.11. It rises shortly after sunset and is prominent all night.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 12h 17m, Decl. 0° 30' N. and transits at 12.42. It is too close to the sun for observation. Conjunction is on the 29th.

Uranus on the 15th is in R.A. 6h 59m, Decl. 23° 07' N. and transits at 7.24.

Neptune on the 15th is in R.A. 13h 10m, Decl. 5° 41' S. and transits at 13.34.

				SEPTEMBER 75th Meridian Civil Time	Min. of Algol	Config. of Jupiter's Sat. 1h 15m
	d	h	m		h m	1
Sat.	1			Annular eclipse of \bigcirc , see p. 29	11 56	0324*
		0	48	σ ξ C ξ 4° 39′ S		
		7	27	σ ♀ Ø ♀ 9° 23′ S		
		7	49	New Moon		
Sun.	2					10234
Mon.	- 3	8	16	σ þ @ þ 4° 43′ N		20143
	•	10		$\sigma \neq \odot$ Inferior		
Tue	4	12	54	$\sigma \Psi \Phi \Psi 5^{\circ} 25' N$	08 45	14023
Wed	5	1.2			00 10	43012
Thur.	6					43120
Fri	7	1		0 Greatest Hell Lat S	05 33	13201
Sat	0	12	16	First Quarter	00 00	4109*
Sat.	0	10	10	B Stationomy in D A		4102
C	0	20				11099
Sun.	10			•••••••••••••••••••••••••••••••••••••••	09 99	49012
mon.	10	15		$\mathbf{M}_{\text{res}} = \mathbf{D}_{\text{res}} + \mathbf{D}_{\text{res}$	02 22	42013
Tue.	11	15		Moon in Perigee. Dist. from \oplus , 228,200 ml	00 11	4103*
wed.	12	8		$\sigma \varphi \varphi \qquad \varphi s^* 21^* N \dots$	23 11	30412
Thur.	13					31204
Fri.	14	2		φ in $\delta\delta$		32014
Sat.	15	7	38	Full Moon	19 59	13024
Sun.	16	11		Greatest elongation W., 17° 53'		dO234
		21	04	o 24 € 24 4° 54′ S		
Mon.	17			*******		20134
Tue.	18	17		۵ in Perihelion	$16 \ 48$	12034
Wed.	19					30124
Thur.	20					31204
Fri.	21				$13 \ 36$	32401
Sat.	22	23	13	C Last Quarter		41302
Sun.	23	2		Q Stationary in R.A		40123
		15	38	\odot enters \simeq , Autumn commences. Long. of \bigcirc 180°		
		16		Moon in Approx. Dist from \oplus 251 200 mi		
Mon	94	10	94	$\sim \alpha	10.95	490.9*
Tue	24	0	24	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	10 20	4203
Tue.	20			•••••••••••••••••••••••••••••••••••••••		42103
wea.	20	1.5	20		07.14	43012
i nur.	27	10	32		07 14	104310
F :	00	23	40	$\begin{array}{ccc} \mathbf{O} \neq \mathbf{U} & \forall \mathbf{I} \uparrow \mathbf{U} \mathbf{I} \uparrow \mathbf{S} \dots \dots \\ \mathbf{B} & \mathbf{C} \qquad \forall \mathbf{I} \uparrow \mathbf{I} \downarrow \mathbf{I} \neq \mathbf{N} \end{array}$		1 20401
rri.	28	23		φ Greatest Hel. Lat. N		32401
Sat.	29	6	0.1	$\sigma P \bigcirc \cdots $		31024
Sun.	30	3	31	σ φ @ φ 3° 46′ Ν	04 02	01324
		20	57	W New Moon		
		21	59	σ 𝒫 Ϣ 🕴 Ϸ 5° 00′ Ν		

Explanation of symbols and abbreviations on p. 4, of time on p. 8.

THE SKY FOR OCTOBER, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During October the sun's R.A. increases from 12h 26m to 14h 22m and its Decl. changes from 2° 47' S. to 14° 06' S. The equation of time changes from +9m 57s to +16m 20s. For changes in the length of the day, see p. 15.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 22. The full moon of October 14th is the Hunter's Moon.

Mercury on the 15th is in R.A. 13h 22m, Decl. 7° 43' S. and transits at 11.53. It is in superior conjunction on the 13th and it is too near the sun all month for observation.

Venus on the 15th is in R.A. 10h 38m, Decl. 5° 48' N. and transits at 9.06. It is a morning star prominent in the eastern sky before sunrise. At this time it has a crescent shape when seen in the telescope.

Mars on the 15th is in R.A. 10h 34m, Decl. 10° 33' N. and transits at 9.02. It is in Leo, passing within a degree of Regulus on the 3rd and moving east of it. It is fainter than Regulus, being of stellar magnitude 1.9.

Jupiter on the 15th is in R.A. 0h 30m, Decl. 1° 31' N. and transits at 22.54 Already risen at sunset, it dominates the sky all night. Opposition is on the 2nd.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 54.

Saturn on the 15th is in R.A. 12h 31m, Decl. 0° 57' S. and transits at 10.57. It is too close to the sun for easy observation.

Uranus on the 15th is in R.A. 7h 01m, Decl. 23° 04' N. and transits at 5.28.

Neptune on the 15th is in R.A. 13h 14m, Decl. 6° 06' S. and transits at 11.40.

				OCTOBER	м	in	Config.
						f	Jupiter's
				75th Meridian Civil Time	Alg	ol	Sat. 23h 45m
	Ь	h	l m		h	m	1
Mon	1	21	30	~ΨΦ Ψ 5° 26' N		111	d2O34
Tue	2	21		~ 20 mist from \oplus 366 800 000 mi			03124
Wod	2	20			00	51	31024
Thur	о Л					01	32014
Thui.	5				91	40	31024
FII.	6	1		$\sim 8 h$ 8 0° 25′ S	21	40	1024
Sat.	7	1 9		$0 \neq 1$ $\psi = 0 = 50 = 5$			40312
Sun.	1			Moon in Perigee. Dist. from \oplus , 229,900 ml	1		42103
		19	00				
3.4	0	19	00	Pirst Quarter	10	•••	49019
won.	ð			· · · · · · · · · · · · · · · · · · ·	10	28	42013
I ue.	9			0 0 1 1 11	1		4032*
Wed.	10	14		Q Greatest brilliancy			43102
Thur.	11	_			15	17	43201
Fri.	12	2		Greatest Hel. Lat. S.			4310*
Sat.	13	5		σΨΟ			40312
		10		σ₿⊙ Superior			
		11		σ ['] [†] Ψ ['] [†] [†] ^{0°} 40′ S			
		23	18	of 24 € 24 4° 36′ S			
Sun.	14	19	51	Full Moon	12	06	12403
Mon.	15						20143
Tue.	16						10324
Wed.	17				08	54	31024
Thur.	18						32014
Fri.	19						31204
Sat.	20	16		Stationary in R.A	05	43	0124*
Sun.	21	12		Moon in Apogee. Dist. from \oplus , 251,300 mi.			12034
		12	10	ර ô € _ ô 3° 52′ S			
Mon.	22			Orionid meteors.			20413
		10		8 in 89			
		18	55	Last Ouarter			
Tue.	23				02	32	41023
Wed.	24					-	d4302
Thur	25				23	21	43201
Fri	26	11	40	~~~?.05′N	20		43120
	20	18	24	$\sim 0^{\circ}$ 0 2° 0 $2^$			10120
Sat	97	10	44				4012*
Sun	41 92	12	22	∠ h € h 5° 99′ N	90	10	44102
Mon	20 20	0	20 19	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	20	10	49019
Tuo	49 20	0	10				41092
iue.	3U 91	ð	04 77		10	-0	41023
wea.	31	0	55	ן ס צ ענ	10	99	30142

THE SKY FOR NOVEMBER, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During November the sun's R.A. increases from 14h 22m to 16h 25m and its Decl. changes from 14° 06' S. to 21° 39' S. The equation of time changes from +16m 20s to a maximum of +16m 23s on the 4th and then to +11m 20s at the end of the month. For changes in the length of the day, see p. 16.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 23.

Mercury on the 15th is in R.A. 16h 31m, Decl. 24° 10' S. and transits at 13.00. It is an evening star but even at greatest eastern elongation on the 28th it is poorly placed for observation.

Venus on the 15th is in R.A. 12h 21m, Decl. 0° 57' S. and transits at 8.48. It is a morning star prominent in the south-eastern sky before sunrise. Greatest western elongation on the 14th. There is a close conjunction with Saturn on the morning of the 21st.

Mars on the 15th is in R.A. 11h 43m, Decl. 3° 33' N. and transits at 8.09. It moves from Leo to Virgo during the month, rising shortly after midnight.

Jupiter on the 15th is in R.A. 0h 19m, Decl. 0° 27' N. and transits at 20.42. It is well up in the east at sunset and sets several hours before dawn. On the 30th it is stationary in R.A. and resumes direct, or eastward, motion among the stars.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 55.

Saturn on the 15th is in R.A. 12h 44m, Decl. 2° 17' S. and transits at 9.08 It is in Virgo west of Spica, rising a few hours before the sun. There is a close conjunction with Venus on the morning of the 21st.

Uranus on the 15th is in R.A. 7h 00m, Decl. 23° 07' N. and transits at 3.25.

Neptune on the 15th is in R.A. 13h 18m, Decl. 6° 30' N. and transits at 9.43.

				NOVEMBER	Min.	Config. of
				75th Meridian Civil Time	Algol	Sat. 22h 00m
	d	h	m		hm	
Thur.	1	17				3204*
Fri.	2	8		Moon in Perigee. Dist. from \oplus , 227,200 mi		32104
		9		o ⁷ Greatest Hel. Lat. N		
		11		ፍ in ይ		
Sat.	3				13 47	30124
Sun.	4					10234
Mon.	5					20134
Tue.	6	1	59	First Quarter	10 36	1034*
Wed.	7					30124
Thur.	8					32104
Fri.	9				07 25	d342O
Sat.	10	0	59	of 2↓ € 2↓ 4° 22′ S		43012
Sun.	11	1				41023
Mon.	12				04 14	42013
Tue.	13	10	52	Full Moon		4103*
Wed.	14	4	-	φ Greatest elongation W., 46° 39'		d4012
Thur.	15				01 03	43120
Fri.	16			Leonid meteors.		32401
Sat	17	19	14	♂ å Ø å 3° 39′ S.	21 52	3042*
Sun	18	8		Moon in Apogee. Dist. from \oplus . 251.900 mi.		10234
Mon	19			······································		20134
Tue	20				18 41	12034
Wed	21	2		9 in Perihelion		03124
wea.	~1	5		$\alpha \circ b = 0^{\circ} 38' S$		00111
		15	01	I Last Quarter	1	
Thur	.99	10	01	8 Greatest Hel Lat S		31204
Fri	22				15 30	32014
Sat	20	6	26	~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~ ~	10 00	3024*
Sun	24	1	59	$\sim h f $ $\sim h 5^{\circ} 51' N$		41032
Sun.	20	19	10	~ 0.6 0 5° 47' N	1	11002
		12	10	\checkmark tu \checkmark tu \checkmark tu \checkmark N		
Mon	96	. 20	19		19 10	42013
Tue	40 97				12 15	41903
Tue.	41	0		B Constant element $E = 91^{\circ} 40'$		40212
wea.	28	0	00	Greatest elongation E., 21 40		40312
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r ri.	30	8	40	Woon in Perigee. Dist. from \oplus , 223,900 mi		45201
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		14		1 '4 Stationary in R.A.		

THE SKY FOR DECEMBER, 1951

Positions of the sun and planets are given for 0h Greenwich Civil Time.

The times of transit at the 75th meridian are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During December the sun's R.A. increases from 16h 25m to 18h 41m and its Decl. changes from 21° 39' S. to 23° 27' S. at the solstice on the 22nd and then to 23° 06' S. at the end of the month. The equation of time changes from +11m 20s to zero on the 25th and then to -3m 00s at the end of the month. For changes in the length of the day, see p. 16.

The Moon—For its phases, perigee and apcgee times and distances, and its conjunctions with the planets, see opposite page. Times of moonrise and moonset are given on p. 23.

Mercury on the 15th is in R.A. 17h 48m, Decl. 22° 11' S. and transits at 12.10. It is in inferior conjunction on the 16th but by the end of the month it is a fairly good morning star which might be seen low in the south-east just before sunrise.

Venus on the 15th is in R.A. 14h 26m, Decl. 11° 49' S. and transits at 8.55. It is a morning star prominent in the south-eastern sky before sunrise.

Mars on the 15th is in R.A. 12h 45m, Decl. 3° 02' S. and transits at 7.13. It is in Virgo just west of Spica, rising shortly after midnight. It has brightened appreciably and ends the year with stellar magnitude 1.3. See Saturn.

Jupiter on the 15th is in R.A. 0h 19m, Decl. 0° 34' N. and transits at 18.44. It is prominently placed in the south-east at sunset and sets shortly after midnight.

For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 55.

Saturn on the 15th is in R.A. 12h 53m, Decl. 3° 12' S. and transits at 7.20. It is in Virgo west of Spica, rising soon after midnight. Mars, moving eastward, overtakes Saturn on the 19th and passes within a degree south of it. At this time Saturn may be distinguished from Mars by its being slightly brighter.

Uranus on the 15th is in R.A. 6h 56m, Decl. 23° 13' N. and transits at 1.23.

Neptune on the 15th is in R.A. 13h 21m, Decl. 6° 48' S. and transits at 7.48.

				DECEMBER	Min.	Config. of
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Thur.	6	7		Q in Perihelion		31024
Fri.	7	5	26	of 2↓ € 2↓ 4° 26′ S	$23 \ 35$	32014
		6		§ Stationary in R.A		
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Sun.	9					0124*
Mon.	10				20 24	2034*
Tue.	11	1		β in Ω		21043
Wed.	12			Geminid meteors		40123
Thur.	13	4	30	Full Moon	17 13	41302
Fri.	14	-		a		43201
Sat	15	0	24	☆ â đ â 3° 32′ S		4310*
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		23		ØGreatest Hel. Lat. N		0.004
Wed.	26					0234*
Thur.	27	0	35	σ ⊈ C		d1024
		3		ξ Stationary in R.A		
		22				
		23		Q Greatest Hel. Lat. N		
Fri.	28	6	43	New Moon	01 19	32041
		18		Moon in Perigee. Dist. from \oplus , 221,900 mi		
Sat.	29					34120
Sun.	30				22 08	43012
Mon.	31			<u> </u>	l	d41O3

PHENOMENA OF JUPITER'S SATELLITES, 1951

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LUNAR OCCULTATIONS

Prepared by J. F. HEARD

When the moon passes between the observer and a star that star is said to be occulted by the moon and the phenomenon is known as a lunar occultation. The passage of the star behind the east limb of the moon is called the immersion and its appearance from behind the west limb the emersion. As in the case of eclipses, the times of immersion and emersion and the duration of the occultation are different for different places on the earth's surface. The tables given below, adapted from the 1951 Nautical Almanac, give the times of immersion or emersion or both for occultations of stars of magnitude 4.5 or brighter visible at Toronto and at Montreal and also at Vancouver and Calgary, at night. Emersions at the bright limb of the moon are given only in the case of stars brighter than magnitude 3.5. The terms *a* and *b* are for determining corrections to the times of the phenomena for stations within 300 miles of the standard stations. Thus if λ_0 , ϕ_0 , be the longitude and latitude of the standard station and λ , ϕ , the longitude and latitude of the neighbouring station then for the neighbouring station we have—

Standard Time of phenomenon = Standard Time of phenomenon at the standard station $+ a(\lambda - \lambda_0) + b(\phi - \phi_0)$

where $\lambda - \lambda_0$ and $\phi - \phi_0$ are expressed in degrees. The quantity P in the table is the position angle of the point of contact on the moon's disk reckoned from the north point towards the east.

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Jan. 17 η Tau 3.0 I 10.1 17 10.4 $ 135$ $Graze$ $$ $$ 17 η Tau 3.0 E 10.1 17 10.4 $ 135$ $Graze$ $$ $$ Feb. 7 Venus -3.3 I 1.6 17 58.8 -1.3 -2.8 107 18 00.4 -1.0 -2 7 Venus -3.3 E 1.6 18 40.7 $+0.3$ $+2.1$ 184 Low $$ $$ $Apr.$ 9 20 Tau 4.0 1 3.5 Sun $$	• • • •
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LUNAR OCCULTATIONS VISIBLE AT TORONTO AND MONTREAL, 1951

LUNAR OCC LTATIONS VISIBLE AT VANCOUVER AND CALGARY, 1951

Dete	Stor	Mag	1	Age	Va	ancouv	er		(Calgary	7	
Date	Stai	mag.	Ē	Moon	P.S.T.	a	b	Р	M.S.T.	a	b	Р
Jan. 11 Feb. 7 13 13-14 May 14 July 30 Aug. 11 18 18 23 23-24	 φ Aqr Venus Q Tau 20 Tau ρ Leo 136 Tau τ Scr λ Aqr λ Aqr 20 Tau 20 Tau 	$\begin{array}{c}$	E IIEIIEEEE	Moon d 4.2 1.6 1.6 7.9 9.2 26.1 9.2 26.1 9.2 15.4 15.4 15.4 21.4 21.4	P.S.1. h m 17 37.6 14 07.1 15 16.8 22 52.1 23 13.0 022 28.5 Low 20 11.3 No occ. No occ. Low 23 08.9	a m -1.6 -0.6 -1.7 -0.6 -0.2 -0.7 -1.6 +0.3	$ \begin{array}{c} $	P 82 20 262 80 108 124 255	M.S.1. h m 18 50.1 15 15 5 16 29.2 23 55.6 0 11.9 23 31.8 2 43.9 Low 1 38.9 1 57.0 0 23 53.0 0 08.8	$\begin{array}{c c} a \\ \hline m \\ -1.4 \\ -0.8 \\ -1.2 \\ -0.5 \\ -0.1 \\ -0.6 \\ +0.5 \\ \hline \\ 0.0 \\ +0.2 \end{array}$	$ \begin{array}{c} $	P 93 32 251 68 96 113 253 338 308 284 249
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METEORS AND METEORITES By Peter M. Millman

A meteor or "shooting star" appears when one of the larger particles comprising the dust of space happens to encounter the earth's atmosphere at high velocity. In general the particle is completely vapourized high in the upper atmosphere but occasionally it is large enough so that a portion reaches the earth's surface, and this solid lump of iron or stone is known as a meteorite. The study of meteors and meteorites contributes a large amount of valuable information concerning the nature and origin of the universe and there are many intriguing problems in this field awaiting solution. The amateur can do work of lasting value here, as the large and very expensive instrumental equipment required for most astronomical research is not needed for the study of meteors.

For any given observation point there is no way of predicting in advance just where the next meteor will appear, in other words, it is chiefly a matter of chance whether it appears north, south, east, west, or directly overhead. Taking an overall average for the whole year and all parts of the night a single observer with an unobstructed view of the sky will see 10 meteors per hour on a clear moonless night. This statement must be qualified by the fact that meteors are roughly twice as numerous during the second half of the night as they are during the first, and their rate of appearance is approximately doubled for the second half of the year as compared with the first six months. There is also a great variation in meteor frequency from one night to the next. The observed meteors range in brightness all the way from those only visible in fairly large telescopes up to great fireballs exceeding the full moon in luminosity. The frequency of meteors increases approximately in inverse proportion to their brightness.

In addition to the stray so-called "sporadic" meteors which appear on any night of the year, there are various swarms of meteors, each swarm moving along in its particular elliptical orbit about the sun. In most cases these meteor orbits are found to correspond closely with those of certain comets. When the earth encounters such a swarm of meteors the apparent paths, when projected backwards in the sky, all seem to meet in a point, a result of perspective. This point indicates the direction from which the meteors are coming and is called the "radiant". The meteor shower is commonly called after the constellation in which the radiant is located. The best known meteor showers are listed in the accompanying table which has been compiled from various sources. Of these showers the Perseids and Geminids are the m st consistent. Some, such as the Leonids, Giacobinids, and Bielids, have provided spectacular displays in certain years and in others have been almost or totally absent. The Bielids have scarcely been observed at all since the 19th century; the Giacobinids were first observed in 1933. The hourly number listed in the table is the approximate number of meteors which are likely to be seen in one hour by a single observer on a clear moonless night at the shower maximum in a normal year.

Amateur cooperation assists greatly in the scientific study of meteors. Visual observations may be divided into two types:

(a) Systematic programs. These may be carried out either by a single observer or by groups of observers. In this case the sky is observed continuously for a period of time and the numbers of meteors seen, their brightness, colour, position, and other characteristics recorded. Plotting the observations on a star map is more important when the program is carried out in cooperation with another party observing some distance away.

(b) The chance observation of a bright meteor or fireball. Any meteor markedly brighter than Jupiter (mag. -2) should be carefully recorded and the observation forwarded to some observatory where meteor records are being kept. In this case it is very important to note the position of the meteor in the sky, as well as

Continued on page 80.

PRINCIPAL ELEMENTS OF THE SOLAR SYSTEM

Planet	Mean D from (a) $\oplus = 1$	Distance Sun) millions of miles	Period (P)	Eccen- tri- city (e)	In- clina- tion (i)	Long. of Node (လူ)	Long. of Peri- helion (π)	Mean Long. of Plan et	
					•	•	•	0	
Mercury	.387	36.0	88.0days	.206	7.0	47.6	76.5	120.5	
Venus.	.723	67.2	224.7	.007	3.4	76.1	130.7	36.0	
Earth	1.000	92.9	365.3	.017			101.9	99.8	
Mars	1.524	141.5	687.0	.093	1.9	49.1	334.9	267.4	
Jupiter	5.203	483.3	11.86yrs.	.048	1.3	99.8	13.3	164.4	
Saturn	9.54	886.	29.46	.056	2.5	113.1	91.8	97.1	
Uranus	19.19	1783.	84.0	.047	0.8	73.7	169.7	76.8	
Neptune	30.07	2793.	164.8	.009	1.8	131.1	44.1	184.0	
Pluto	39.46	3666.	247.7	.249	17.1	109.5	223.4	158.3	

ORBITAL ELEMENTS (1944, Dec. 31, 12^h)

PHYSICAL ELEMENTS

Object	Symbol	Mean Dia- meter miles	Mass $\oplus = 1$	Density water = 1	Axial Rotation	Mean Sur- face Grav- ity $\bigoplus = 1$	Albedo Bond's	Ma tuc Op tio Elo ti	agni- le at posi- on or onga- ion
Sun	Ċ	864,000	332,000	1.4	24 ^d 7 (equa-	27.9		_	26.7
Moon	Q	2,160	.0123	3.3	$27^{d} 7.7^{h}$. 16	.07	_	12.6
Mercury	Ę	3,010	.056	3.8	88 ^d	.27	.07		$0\pm$
Venus	Q	7,580	.82	4.9	30 ^d ?	.85	.59	_	$4\pm$
Earth	\oplus	7,918	1.00	5.5	$23^{h} 56^{m}$	1.00	.29		
Mars	5	4,220	.108	4.0	$24^{\rm h}$ $37^{\rm m}$.38	.15	_	$2\pm$
Jupiter	24	87,000	318.	1.3	$9^{h} 50^{m} \pm$	2.6	.56?	_	$2\pm$
Saturn	b	72,000	95.	.7	$10^{b}15^{m} \pm$	1.2	.63?		$0\pm$
Uranus	ð	31,000	14.6	1.3	$10^{h}.8\pm$.9	.63?	+	5.7
Neptune	W	33,000	17.2	1.3	16 ^h ?	1.0	.73?	+	7.6
Pluto	P	4,000?	.8 ?					÷	14

SATELLITES OF THE SOLAR SYSTEM

SATELLITE OF THE EARTH Miles Miles Miles SATELLITE OF THE EARTH Moon $ -12.6 $ 530 238,857 27 07 43 2160 SATELLITES OF MARS Phobos 12 8 5,800 0 07 39 10? Hall, 1877 Deimos 13 21 14,600 1 6 18 57 Hall, 1877 SATELLITES OF JUPITER V 13 48 112,600 0 11 57 100? Barnard, 1892 Io 5 112 261,800 18 282,000 Galileo, 1610 Callisto 6 178 416,600 3 13 14 2000 Galileo, 1610 Callisto 6 499 1,60,000 16 63 2200 Galileo, 1610 Callisto 6 173 7,292,000/260 157 Nicholson, 1938 X1 18 5990 14,000,000 f92 157 Nicholson, 1938 X1 18<	Name	Stellar Mag.	Mean F	Dist. from Planet	Re	volut Perio h	tion d m	Diamete Miles	r Discoverer
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Moon $ -12.6 $ 530 $ 238,857 $ 27 07 43 2160 SATELLITES OF MARS Phobos 12 8 5,800 0 07 39 10? Hall, 1877 Deimos 13 21 14,600 1 06 18 5? Hall, 1877 SATELLITES OF JUPITER V 13 48 112,2600 0 11 57 100? Barnard, 1892 Ganymede 5 284 664,200 7 03 43 2000 Galileo, 1610 Callisto 6 499 1,169,000 16 16 32 3200 Galileo, 1610 Callisto 6 499 1,400,0250 16 100? Perrine, 1904 102? Perrine, 1904 11 16 313 7,292,000260 15? Nicholson, 1938 Nitholson, 1938 15? Nicholson, 1938 15? Nicholson, 1914 SATELLITES OF SATURN Mimas 12 27 115,000 <td>SATELLITE</td> <td>of the]</td> <td>Earth</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td>	SATELLITE	of the]	Earth						
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Phobos 12 8 5,800 0 07 39 10? Hall, 1877 Deimos 13 21 14,600 1 06 18 57 Hall, 1877 SATELLITES OF JUPITER V 13 48 112,600 0 11 57 100? Barnard, 1892 Io 5 112 261,800 1 18 28 2:00 Galileo, 1610 Ganymede 5 284 664,200 7 03 3200 Galileo, 1610 Callisto 6 499 1,169,000 16 632 3200 Galileo, 1610 Callisto 6 499 1,70,000250 16 100? Perrine, 1904 VII 14 3037 7,114,000,250 16 100? Molson, 1938 X1 18 5990 14,000,000 22 37 Molson, 1938 VIII 16 6240 14,600,000 23 3500? W. Herschel, 1789 <td>SATELLITES</td> <td>of Ma</td> <td>RS</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td>	SATELLITES	of Ma	RS						
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VII10 0.240 $14,000,000$ 7.59 401Inference, 1906IX17 6360 $14,900,000$ 7.58 20?Nicholson, 1914SATELLITES OF SATURNMimas1227 $115,000$ 02237400?W. Herschel, 1789Enceladus1234 $148,000$ 10853500?W. Herschel, 1789Tethys1143 $183,000$ 12118800?G. Cassini, 1684Dione1155 $234,000$ 21741700?G. Cassini, 1684Rhea1076 $327,000$ 412251100?G. Cassini, 1672Titan8177759,0001522412600?Huygens, 1655Hvperion13214920,000210638300?G. Cassini, 1671Phoebe141870 $8,034,000$ 550200?W. Pickering, 1898SATELLITES OF URANUSMiranda179 $81,000$ 10956Kuiper, 1948Lassell1614119,00021229600?Lassell, 1851Umbriel1619166,00040328400?Lassell, 1851Umbriel1619166,000131107900?W. Herschel, 1787Oberon1442364,000131107900?W. Herschel, 1787<		18	0990 6940	14,000,000	720			101	Melette 1009
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Rhea1076327,000412251100?G. Cassini, 1672Titan8177759,0001522412600?Huygens, 1655Hvperion13214920,000210638300?G. Bond, 1848Iapetus115152,210,0007907561000?G. Cassini, 1671Phoebe1418708,034,000550200?W. Pickering, 1898SATELLITES OFURANUSMiranda17981,00010956Kuiper, 1948Ariel1614119,00021229600?Lassell, 1851Umbriel1619166,00040328400?Lassell, 1851Titania1432272,000816561000?W. Herschel, 1787Oberon1442364,000131107900?W. Herschel, 1787SATELLITE OFNEPTUNETriton1316220,000521033000?Lassell, 1846Nereid194005,500,0002yr.200?Kuiper, 1949	Dione	11	55	234,000	2	17	41	700?	G. Cassini, 1684
Titan 8 177 759,000 15 22 41 2600? Huygens, 1655 Hvperion 13 214 920,000 21 06 38 300? G. Bond, 1848 Iapetus 11 515 2,210,000 79 07 56 1000? G. Cassini, 1671 Phoebe 14 1870 8,034,000 550 200? W. Pickering, 1898 SATELLITES OF URANUS 200? W. Pickering, 1898 Miranda 17 9 81,000 1 09 56 Kuiper, 1948 Ariel 16 14 119,000 2 12 29 600? Lassell, 1851 Umbriel 16 19 166,000 4 03 28 400? Lassell, 1851 Titania 14 32 272,000 8 16 56 1000? W. Herschel, 1787 Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 <td>Rhea</td> <td>10</td> <td>76</td> <td>327,000</td> <td>4</td> <td>12</td> <td>25</td> <td>1100?</td> <td>G. Cassini, 1672</td>	Rhea	10	76	327,000	4	12	25	1100?	G. Cassini, 1672
Hyperion13214920,0002106383007G. Bond, 1848Iapetus115152,210,0007907561000?G. Cassini, 1671Phoebe1418708,034,00050200?W. Pickering, 1898SATELLITES OFURANUSMiranda17981,00010956Kuiper, 1948Ariel1614119,00021229600?Lassell, 1851Umbriel1619166,00040328400?Lassell, 1851Titania1432272,000816561000?W. Herschel, 1787Oberon1442364,000131107900?W. Herschel, 1787SATELLITE OFNEPTUNETriton1316220,000521033000?Lassell, 1846Nereid194005,500,0002yr.200?Kuiper, 1949	Titan	8	177	759,000	15	22	41	2600?	Huygens, 1655
Tapetus115152,210,00079075010007G. Cassini, 1671Phoebe141870 $8,034,000$ 550 200?W. Pickering, 1898SATELLITES OFURANUSMiranda179 $81,000$ 10956Kuiper, 1948Ariel1614119,00021229600?Lassell, 1851Umbriel1619166,00040328400?Lassell, 1851Titania1432272,000816561000?W. Herschel, 1787Oberon1442364,000131107900?W. Herschel, 1787SATELLITE OFNEPTUNETriton1316220,000521033000?Lassell, 1846Nereid194005,500,0002yr.200?Kuiper, 1949	Hyperion		214	920,000	21	06	38	300?	G. Bond, 1848
SATELLITES OF URANUS Kuiper, 1948 Miranda 17 9 $81,000$ 1 09 56 Kuiper, 1948 Ariel 16 14 119,000 2 12 29 600? Lassell, 1851 Umbriel 16 19 166,000 4 03 28 400? Lassell, 1851 Titania 14 32 272,000 8 16 1000? W. Herschel, 1787 Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Iapetus Phosho		010 1970	2,210,000	19	07	50	10001	G. Cassini, 10/1
SATELLITES OF URANUS Miranda 17 9 $81,000$ 1 09 56 Kuiper, 1948 Ariel 16 14 119,000 2 12 29 600? Lassell, 1851 Umbriel 16 19 166,000 4 03 28 400? Lassell, 1851 Titania 14 32 272,000 8 16 1000? W. Herschel, 1787 Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Phoebe	14	1870	0,034,000	000		1	2001	w. Pickering, 1898
Miranda 17 9 81,000 1 09 56 Kuiper, 1948 Ariel 16 14 119,000 2 12 29 600? Lassell, 1851 Umbriel 16 19 166,000 4 03 28 400? Lassell, 1851 Titania 14 32 272,000 8 16 56 1000? W. Herschel, 1787 Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	SATELLITES	OF UR	ANUS						
Ariel 16 14 119,000 2 12 29 600? Lassell, 1851 Umbriel 16 19 166,000 4 03 28 400? Lassell, 1851 Titania 14 32 272,000 8 16 56 1000? W. Herschel, 1787 Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Miranda	17	9	81,000	1	09	561		Kuiper, 1948
Umbriel 16 19 166,000 4 03 28 400? Lassell, 1851 Titania 14 32 272,000 8 16 56 1000? W. Herschel, 1787 Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Ariel	16	14	119,000	2	12	29	600?	Lassell, 1851
Titania 14 32 272,000 8 16 56 1000? W. Herschel, 1787 Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Umbriel	16	19	166,000	4	03	28	400?	Lassell, 1851
Oberon 14 42 364,000 13 11 07 900? W. Herschel, 1787 SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Titania	14	32	272,000	8	16	56	1000?	W. Herschel, 1787
SATELLITE OF NEPTUNE Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Oberon	14	42	364,000	13	11	07	900?	W. Herschel, 1787
Triton 13 16 220,000 5 21 03 3000? Lassell, 1846 Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	SATELLITE	of Nep	TUNE						
Nereid 19 400 5,500,000 2 yr. 200? Kuiper, 1949	Triton	13	16	220,000	5	21	03	3000?	Lassell, 1846
	Nereid	19	400	5,500,000	-	2 yr.		200?	Kuiper, 1949

*As seen from the sun.

Satellites Io, Europa, Ganymede, Callisto are usually denoted I, II, III, IV respectively, in order of distance from the planet.

Much pleasure may be derived from the estimation of the brightness of variable stars. Maps of the fields of four bright variable stars are given below. In each case the magnitudes of several suitable comparison stars are given. These magnitudes are given as magnitudes, tenths and hundredths, with the decimal point omitted. Thus a star 362 is of magnitude 3.62. To determine the brightness of the variable at any time, carefully estimate the brightness as some fraction of the interval between two comparison stars, one brighter and one fainter than the variable. The result may then be expressed in magnitudes and tenths. Record the magnitude and time of observation. When a number of observations have been made, a graph may be plotted showing the magnitude estimate as ordinates against the date (days and tenths of a day) as abscissae. Such studies of naked-eye estimates of brightness will at once reveal the differences in variation between the different kinds of variable. For each short period variable the observations made on any one cycle may be carried forward one, two or any number of periods to form a combined light curve.

For the two cepheids, good mean curves may be readily found by observing the variables once a night on as many nights as possible. For Algol, which changes rapidly for a few hours before and after minimum, estimates should be made at quarter or half hour intervals around the times of minimum as tabulated on pages 31-53. Mira may be observed for a couple of months as it rises from the naked-eye limit to 2nd or 3rd magnitude maximum and fades again.



N	lame	Design.	Max.	Min.	Sp.	Period	Type	Date	Discoverer
η Ν εδ U	Aql Aql Aur Cep Cep	$\begin{array}{r} 194700 \\ 184300 \\ 045443 \\ 222557 \\ 005381 \end{array}$	$ \begin{array}{r} 3.7 \\ -0.2 \\ 3.3 \\ 3.6 \\ 6.8 \end{array} $	$\begin{array}{r} 4.4 \\ 10.9 \\ 4.1 \\ 4.3 \\ 9.2 \end{array}$	G4 Q F5p G0 A0	7.17652 Irr. 9833. 5.36640 2.49293	Cep Nova Ecl Cep Ecl	$ 1784 \\ 1918 \\ 1821 \\ 1784 \\ 1880 $	Pigott Bower Fritsch Goodricke W. Ceraski
o RR R X P	Cet ¹ Cet CrB Cyg Cyg	021403 012700 154428 194632 201437a	$2.0 \\ 8.4 \\ 5.8 \\ 4.2 \\ 3.5$	$10.1 \\ 9.0 \\ 13.8 \\ 14.0 \\ 6.0$	M5e F0 cG0e M7e B1qk	331.8 0.55304 Irr. 412.9 Irr.	LPV Clus RCrB LPV Nova	$1596 \\ 1906 \\ 1795 \\ 1686 \\ 1600$	Fabricius Oppolzer Pigott Kirch Blaeu
SS XX Ç N R	Cyg Cyg Gem Gem Gem	213843 200158 065820 060822 070122a	$8.1 \\ 11.4 \\ 3.7 \\ 3.3 \\ 6.5$	$12.0 \\ 12.1 \\ 4.1 \\ 4.2 \\ 14.3$	Pec. A cG1 M2 Se	Irr. 0.13486 10.15353 235.58 370.1	SSCyg Clus Cep LPV LPV	1896 1904 1847 1865 1848	Wells L. Ceraski Schmidt Schmidt Hind
U a R B	Gem Her Hya Leo Lyr	$\begin{array}{c} 074922 \\ 171014 \\ 132422 \\ 094211 \\ 184633 \end{array}$	$8.8 \\ 3.1 \\ 3.5 \\ 5.0 \\ 3.4$	$13.8 \\ 3.9 \\ 10.1 \\ 10.5 \\ 4.3$	Pec. M5 M7e M7e B5e	Irr. Irr. 414.7 310.3 12.92504	SSCyg SemiR LPV LPV Ecl	1855 1795 1670 1782 1784	Hind W. Herschel Montanari Koch Goodricke
RR а U В р	Lyr Ori² Ori Per³ Per	$\begin{array}{c} 192242\\ 054907\\ 054920\\ 030140\\ 025838\end{array}$	$7.2 \\ 0.2 \\ 5.4 \\ 2.3 \\ 3.3$	$\begin{array}{r} 8.0 \\ 1.2 \\ 12.2 \\ 3.5 \\ 4.1 \end{array}$	A5 M2 M7e B8 M4	0.56685 2070.Irr. 376.9 2.86731 Irr.	Clus SemiR LPV Ecl Irr.	$ 1901 \\ 1840 \\ 1885 \\ 1669 \\ 1854 $	Fleming J. Herschel Gore Montanari Schmidt
R R λ RV SU	Sge Sct Tau Tau Tau	200916 1842 <i>o</i> 5 035512 044126 054319	$8.6 \\ 4.5 \\ 3.8 \\ 9.4 \\ 9.5$	$10.4 \\ 9.0 \\ 4.1 \\ 12.5 \\ 15.4$	cG7 K5e B3 K0 G0e	70.84 141.5 3.95294 78.60 Irr.	SemiR SemiR Ecl SemiR RCrB	1859 1795 1848 1905 1908	Baxendell Pigott Baxendell L. Ceraski Cannon
a N N	UMi⁴ Her Lac	$\begin{array}{c} 012288 \\ 180445 \\ 221255 \end{array}$	$2.3 \\ 1.5 \\ 2.2$	$2.4 \\ 14.0 \\ -$	cF7 Q Q	3.96858 Irr. Irr.	Cep Nova Nova	1911 1934 1936	Hertzsprung Prentice Peltier

REPRESENTATIVE BRIGHT VARIABLE STARS

¹oCet (Mira); ²aOri (Betelgeuse); ³βPer (Algol); ⁴aUMi (Polaris).

The designation (Harvard) gives the 1900 position of the variable; here the first two figures give the hours, and the next two figures the minutes of R.A., while the last two figures give the declination in degrees, italicised for southern declinations. Thus the position of the fourth star of the list, δ Cep (222557) is R.A. 22h 25m, Dec. + 57°. The period is in days and decimals of a day. The type is based on the classification of Gaposchkin and Gaposchkin's comprehensive text-book, *Variable Stars*. The abbreviations here used are: Ecl, Eclipsing Binaries; LPV, Long Period Variables; Semi R, Semiregular; Cep, Cepheids; Clus, cluster type; Nova; SS Cyg and R Cr B, irregular variables of which SS Cygni and R Coronae Borealis are prototypes; and Irr, other irregular variables.

DOUBLE AND MULTIPLE STARS

A number of the stars which appear as single to the unaided eye may be separated into two or more components by field glasses or a small telescope. Such objects are spoken of as *double* or *multiple stars*. With larger telescopes pairs which are still closer together may be resolved, and it is found that, up to the limits of modern telescopes, over ten per cent. of all the stars down to the ninth magnitude are members of double stars.

The possibility of resolving a double star of any given separation depends on the diameter of the telescope objective. Dawes' simple formula for this relation is d'' = 4.5/A, where d is the separation, in seconds of arc, of a double star that can be just resolved, and A is the diameter of the objective in inches. Thus a one-inch telescope should resolve a double star with a distance of 4".5 between its components, while a ten-inch telescope should resolve a pair 0".45 apart. It should be noted that this applies only to stars of comparable brightness. If one star is markedly brighter than its companion, the glare from the brighter makes it impossible to separate stars as close as the formula indicates. This formula may be applied to the observation of double stars to test the quality of the seeing and telescope.

It is obvious that a star may appear double in one of two ways. If the components are at quite different distances from the observer, and merely appear close together in the sky the stars form an *optical* double. If, however, they are in the same region of space, and have common proper motion, or orbital motion about one another, they form a *physical* double. An examination of the probability of stars being situated sufficiently close together in the sky to appear as double shows immediately that almost all double stars must be physical rather than optical.

Double stars which show orbital motion are of great astrophysical importance, in that a careful determination of their elliptical orbits and parallaxes furnishes a measure of the gravitational attraction between the two components, and hence the mass of the system.

In the case of many unresolvable close doubles, the orbital motion may be determined by means of the spectroscope. In still other doubles, the observer is situated in the orbital plane of the binary, and the orbital motion is shown by the fluctuations in light due to the periodic eclipsing of the components. Such doubles are designated as *spectroscopic* binaries and *eclipsing* variables.

The accompanying table provides a list of double stars, selected on account of their brightness, suitability for small telescopes, or particular astrophysical interest. The data are taken chiefly from Aitken's New General Catalogue of Double Stars, and from the Yale Catalogue of Bright Stars. Successive columns give the star, its 1950 equatorial coordinates, the magnitudes and spectral classes of its components, their separation, in seconds of arc, and the approximate distance of the double star in light years. The last column gives, for binary stars of well determined orbits, the period in years, and the mean separation of the components in astronomical units. For stars sufficiently bright to show colour differences in the telescope used, the spectral classes furnish an indication of the colour. Thus O and B stars are bluish white, A and F white, G yellow, K orange and M stars reddish.

A good reference work in the historical, general, and mathematical study of double stars is Aitken's *The Binary Stars*.

REPRESENTATIVE DOUBLE STAF

9	Star		a 19	50 δ		Mag.	and Spect.	d	D	Remarks
π η α γ a	And Cas UMi Ari Pis	h 00 00 01 01 01	m 34 2 46.0 48.8 50.8 59.4	$^{\circ}$ +33 +57 +89 +19 +02	27 33 02 03 31	4.4B3; 8 3.6F8; 7 var. F8; 4.8A0; 4 5.2A2; 4	.5 .2M0 8.8 .8A0 .3A2	"36 8 19 8 3 2.4	L.Y. 470 18 470 150 130	† 526y; 66AU Polaris ††
$egin{array}{c} \gamma \\ 6 \\ \eta \\ 32 \\ \beta \end{array}$	And Tri Per Eri Ori	02 02 02 03 05	00 8 09.5 47.0 51.8 12.1	+42 + 30 + 55 - 03 - 08	05 04 41 06 15	2 3K0; 5 5.4G4; 7 3.9K0; 8 5.0A; 6. 0.3B8; 7	5.4A0; 6.6 COF3 3.5 3G5 .0	$10, 0.7 \\ 3.6 \\ 28 \\ 6.7 \\ 9$	410 330 540 300 540	56y; 23AU tt
θ β 12 α δ	Ori Mon Lyn CMa Gem	05 06 06 06 07	$32.8 \\ 26.4 \\ 41.8 \\ 43.0 \\ 17.1$	$-05 \\ -07 \\ +59 \\ -16 \\ +22$	25 00 30 39 05	5.4;6.8; 4.7B2; 5 5.3A2; 6 -1.6A0; 3.5F0; 8	6.8; 7.9; O .2; 5.6 .2; 7.4 ; 8.5F .0M0	$13, 17 \\7, 25 \\1.7, 8 \\11 \\6.8$	540 470 180 9 58	Trapezium † 50y; 20AU †
αζγξι	Gem Cnc Leo UMa Leo	07 08 10 11 11	$31.4 \\ 09.3 \\ 17.2 \\ 15.5 \\ 21.3$	+32 + 17 + 20 + 31 + 10	00 48 06 48 48	2.0A0; 2 5.6G0; 6 2.6K0; 3 4.4G0; 4 4.1F3; 6	.8A0; 9M10 5.0; 6.2 5.8G5 9G0 .8F3	${}^{4, 70}_{1, 5}_{4}_{2}_{2}$	47 78 160 25 69	340y; 79AU 60y; 21AU 400y ††60y; 20AU
γαζπε	Vir CVn UMa Boo Boo	12 12 13 14 14	$39.1 \\ 53.7 \\ 21.9 \\ 38.4 \\ 42.8$	-01 + 38 + 55 + 16 + 27	10 35 11 38 17	3.6F0;3 2.9A0;5 2.4A2;4 4.9A0;5 2.7K0;5	.7F0 .4A0 .0A2 .1A0 5.1A0	$ \begin{array}{r} 6 \\ 20 \\ 14 \\ 6 \\ 3 \end{array} $	34 140 78 360 220	171y; 42AU †† †† †
w0w000	Boo Ser Sco Her Her	14 15 16 17 17	$\begin{array}{r} 49.1 \\ 32.4 \\ 01.6 \\ 12.4 \\ 13.0 \end{array}$	$^{+19}_{+10}_{-11}_{+14}_{+24}$	18 42 14 27 54	4.8G5; 6 4.2F0; 5 5.1F3; 4 var.M5; 3.2A0; 8	9.7 .2F0 .8; 7G7 5.4G .1G2	$3 \\ 4 \\ 1, 7 \\ 5 \\ 11$	22 170 84 540 100	151y; 31AU 44.7y; 19AU † † Optical
έ β α γ 61	Lyr Cyg Cap Del Cyg	18 19 20 20 21	$\begin{array}{r} 42.7 \\ 28.7 \\ 14.9 \\ 44.3 \\ 04.6 \end{array}$	$+39 \\ +27 \\ -12 \\ +15 \\ +38$	37 51 40 57 30	5.1, 6.0. 3.2K0; 5 3.8G5; 4 4.5G5; 5 5.6K5; 6	A3; 5.1 , 5.4A 5 5.4B9 6G0 5.5F8 5.3K5	3, 2 34 376 10 23	200 410 110 11	Pairs 207" † Optical
β5608 6	Cep Aqr Cep Lac Cas	21 22 22 22 23	$\begin{array}{r} 28.1 \\ 26.2 \\ 27.3 \\ 33.6 \\ 56.5 \end{array}$	$+70 \\ -00 \\ +58 \\ +39 \\ +55$	20 17 10 23 29	var.B1; 8 4.4F2; 4 var.G0; 5 5.8B3; 6 5.1B2; 7	8.0A3 .6F1 7.5A0 .5B5 .2B3	14 3 41 22 3	540 140 650 1100 820	t t

t or tt, one, or two of the components are themselves very close visual double or, more generally, spectroscopic binaries.

THE BRIGHTEST STARS†

Their Magnitudes, Types, Proper Motions, Distances and Radial Velocities

The accompanying table contains the principal facts regarding 259 stars brighter than apparent magnitude 3.51 which it is thought may be of interest to our amateur members. The various columns should be self-explanatory but some comments may be in order.

The first column gives the name of the star and if it is preceded by the sign || such means that the star is a visual double and the combined magnitude is entered in the fourth column. Besides the 48 thus indicated there are 12 others on the list with faint companions but for these it is not thought that there is any physical connection. In the case of the 20 stars variable in light this fourth column shows their maximum and minimum magnitudes. The 19 first magnitude stars are set up in bold face type.

In the fifth column are given the types as revised at various observatories principally at our own, but omitting the s and n designations descriptive of the line character. The annual proper motion follows in the next column and this may not necessarily be correct to the third decimal place.

The parallaxes are taken from the Yale Catalogue of Stellar Parallaxes 1935, the mean of the trigonometric and spectroscopic being adopted. The few negative trigonometric parallaxes were adjusted by Dyson's tables before being combined with the spectroscopic. The distance is given also in light years in the eighth column as to the lay mind that seems a fitting unit. The absolute magnitudes in the ninth column are the magnitudes the stars would have if all were at a uniform distance of 32.6 light years ($\pi = 0$."1). At that distance the sun would appear as a star of magnitude 4.8.

The radial velocities in the last column have been taken from Vol. 18 of the Lick Publications. An asterisk * following the velocity means that such is variable. In these cases the velocity of the system, if known, is given; otherwise a mean velocity for the observations to date is set down.

Of the 259 stars or star systems here listed 146 are south and 113 north of the equator. This is to be expected from the fact that the northern half of the sky includes less of the Milky Way than the southern.

The number in each spectral class, apart from the one marked peculiar, is as follows: O, 3: B, 74; A, 55; F, 22; G, 43, K, 42 and M, 19. The B-stars are intrinsically luminous and appear in this list out of all proportion to their total number. The stars in Classes A and K are by far the most numerous but the revision of types throws many originally labelled K back into the G group.

From the last column we see that 98 velocities are starred, indicating that 38 per cent of the bright stars, or at least one in every three, are binary in character. For visual binaries the proportion has usually been listed as one in nine. Our list shows one in six but it is only natural to expect that we would observe a higher proportion among the nearby stars, such as these are on the average.

Other relationships can be established from the list if our amateur members care to study it.

[†]This feature of the HANDBOOK, first appearing in the 1925 edition, was prepared and frequently revised by the late Dr. W. E. Harper (1878-1940).

and some state of the second									
Star	R.A. 1950	Decl. 1950	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
a Andr β Cass γ Pegs β Hydi a Phoe δ Andr a Cass β Ceti γ Cass	h m 0 6 11 23 24 37 38 41 54	$\begin{array}{c} & & & \\ & & & \\ +28 & 49 \\ +58 & 52 \\ +14 & 54 \\ -77 & 32 \\ -42 & 35 \\ +30 & 35 \\ +30 & 35 \\ +56 & 16 \\ -18 & 16 \\ +60 & 27 \end{array}$	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	A1 F2 B2 G0 G5 K3 G8 G7 B0e	.217 .561 .015 2.243 .448 .167 .062 .233 .031	.034 .080 .005 .162 .040 .026 .018 .052 .035	96 41 652 21 81 125 181 63 93	$ \begin{array}{c} -0 & 1 \\ 1 & 9 \\ -3 & 6 \\ 4 & 0 \\ 0 & 4 \\ 0 & 6 \\ -1 & 5 \\ 0 & 8 \\ -0 & 1 \end{array} $	$\begin{array}{c} \textbf{km}./\textbf{sec.}\\ -13.0^{*}\\ +11.4\\ +5.0^{*}\\ +22.8\\ +74.6^{*}\\ -7.1^{*}\\ -3.8\\ +13.1\\ -6.8\end{array}$
$ \begin{aligned} & \beta \text{ Phoe}, \\ &\beta \text{ Andr}, \\ &\delta \text{ Cass}, \\ &\gamma \text{ Phoe}, \\ &\mathbf{a} \text{ Erid}, \\ &\mathbf{a} \text{ Erid}, \\ &\mathbf{a} \text{ Cass}, \\ &\beta \text{ Arie}, \\ &\mathbf{a} \text{ Hydi}, \end{aligned} $	1 04 07 23 26 36 49 51 52 57	$\begin{array}{r} -46 & 59 \\ +35 & 21 \\ +59 & 59 \\ -43 & 34 \\ -57 & 29 \\ +89 & 02 \\ +63 & 25 \\ +20 & 34 \\ -61 & 49 \end{array}$	$\begin{array}{r} 3.4\\ 2.4\\ 2.8-2.9\\ 3.4\\ 0.6\\ 2.3-2.4\\ 3.4\\ 2.7\\ 3.0\end{array}$	G4 M0 A3 M1 B9 F7 B5 A3 A7	.043 219 .308 223 .093 .043 .043 .150 .255	.020 .041 .050 .008 .046 .008 .011 .066 .080	163 79 65 407 71 407 296 49 41	$ \begin{array}{r} -0.1 \\ 0.5 \\ 1.3 \\ -2.1 \\ -1.1 \\ -3.4 \\ -1.4 \\ 1.8 \\ 2.5 \\ \end{array} $	$ \begin{array}{r} -1.2 \\ +0.1 \\ +6.8 \\ +25.7^* \\ +19 \\ -17.4^* \\ -8.1 \\ -0.6^* \\ +7.0^* \end{array} $
γ Andr a Arie β Tria o Ceti θ Erid	2 01 04 07 17 56	+42 05 +23 14 +34 45 - 3 12 -40 30	2.3 2.2 3.1 1.7-9.6 3.4	K0 K2 A6 M6e A2	.073 .242 .161 .239 068	. 020 045 029 013 . 032	163 72 112 251 102	$ \begin{array}{r} -1 & 2 \\ 0 & 5 \\ 0 & 4 \\ -2 & 7 \\ 0 & 9 \end{array} $	-11 7-14.3+10.4*+57.8*+11.9*
a Ceti γ Pers β Pers δ Pers δ Pers γ Hydi γ Hydi γ Ers γ Ers γ Erid λ Taur	3 00 01 02 05 21 39 45 48 51 54 56 58	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{c} 2.8\\ 3.1\\ 3.3-4.1\\ 2.1-32\\ 1.9\\ 3.1\\ 3.0\\ 3.2\\ 2.9\\ 3.0\\ 3.2\\ 3.8-42\end{array}$	M1 F9 M6 B8 F4 B5 B5p M3 B1 B2 M0 B3	080 .012 .176 011 .041 .047 053 .124 .023 .041 .133 015	018 017 024 033 017 012 014 008 008 006 012 008	181 192 136 99 192 272 233 407 407 543 272 407	$ \begin{array}{c} -0 & 9 \\ -0 & 7 \\ 0 & 3 \\ -2 & 0 \\ -1 & 5 \\ -1 & 3 \\ -2 & 3 \\ -2 & 6 \\ -3 & 1 \\ -1 & 6 \\ -2 & 2 \end{array} $	$\begin{array}{r} -25.7 \\ + 1.0^{*} \\ +28.2 \\ + 5.7^{*} \\ - 2.4 \\ -10.^{*} \\ +10.3 \\ +16.0 \\ +20.9 \\ - 6 \\ +61.7 \\ +13.0^{*} \end{array}$
a Reti	4 14	-62 36	3.4	G5	070	.016	204	-0 6	+35.6

aU. Min., Polaris: R.A. 1h 49.5m; Dec. + 89° 02' (1951)

Star	R.A. 1950	Decl. 1950	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
	h m	0 /		VQ	, ,, ,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,	// 060	54	0.0	km /sec.
a Dam	4 00	710 24		NO AO	. 205	.000	- 04	0.0	+04.1
	33	-55 09	0.0	AUD		1.1.24			+20.0
#• Orio	41	+ 0.52	3.3	F D IZA	.4/4	.124	20	3.8	+24.0
4 Auri	54	+33 05	2.9	K4	.030	.020	163	-0.0	+11.0
€ Auri	58	+43 45	3.1-3.8	F2	.015	.006	543	-2.7	-4.1 *
η Auri	5 03	+41 10	3.3	B3	.082	013	251	-1.1	+7.8
€ Leps	03	-22 26	3.3	K5	.074	.016	204	-0.7	+1.0
<i>B</i> Erid	05	- 5 09	2.9	A1	.117	055	59	1.6	- 7
μ Leps	11	-16 16	3.3	A0p	.053	.020	163	-0.2	+27.7
β Orio	12	- 8 15	0.3	B8p	.005	006	543	-5.8	$+23.6^{*}$
lla Auri	13	+4557	0.2	G1	.439	078	42	-0.3	+30.2
η Orio	22	- 2 26	3.4	B0	.009	.006	543	-2.7	$+19.5^{*}$
γ Orio	22	+ 6 18	1.7	B2	.019	015	217	-2.4	+18.0
<i>B</i> Taur	23	+28 34	1.8	B8	.180	028	116	-1.0	+ 8.0
β Lens	26	-2048	3.0	G2	095	018	181	-0.7	-13.5
δ Orio	29	-0.20	24-25	BO	.006	007	466	-3.4	+19.9*
a Lens	31	-1751	2 7	F6	006	012	272	-21	± 24.7
4 Orio	33	- 5 56	2.0	08	007	021	155	-0.5	+21 5*
e Orio	34	- 1 14	1.8	B0	004	008	407	-3.7	± 25.8
t Taur	35	+2107	3.0	B3e	028	010	326	-2.0	$\pm 16.4^{*}$
Il Orio	38	- 1 58	1.8	BO	012	011	206	-3.0	+18.8
	38	-34 06	2.0	BS	036	022	148	_0.6	134 6
	45	_ 0 41	2.0	BO	.000	006	543	-30	+90.1
R Colm	40	-35 47	2.2	KU	307	026	195	0.3	+20.1
a Orio	59	1 7 94	0 5 1 1	MO	022	012	120	4 1	1 21 0*
R Auri	56		0.0-1.1	A0p	016	052	62	07	18 1*
	56	± 27.12	2.1-2.2		1040	0.002	110	0.7	-10.1
10 Auii	00	T 37 13	2.1	Л	. 100	.029	112	0.0	+20.0
n Gemi	6 12	+22.31	3 2-4 2	М2	062	014	233	-1.1	+21.4*
ζ C Mai	18	-30.02	3 1	B3	012	013	251	-0.7	+33.1*
u Cemi	20	$\pm 22 32$	3.1	MS	120	016	201	_0.9	1 54 8
A C Mai	20	-1756	20	R1	. 129	014	207	-0.0	1 34 4*
p C maj	20	-52 40	0	FO	000	005	652	-7 4	+ 20 5
v Gemi	20	16 97	1 0	42	066	050	65	0 4	11 2*
	26	_43 00	29	Rg	021	000	142	0.4	1 28 2*
• Cemi	1 1	195 19	3.2	C0	021	0020	362		+ 0.0
Ł Gemi	40	119 57	0.4	50	.020	054	60	- 2.0	+ 0.0
	42	-16 20	0.4	го А9	1 215	2004	00	1 2	7 5*
a Dict	40	-10 39	2.0	A5	1.010	. 300	0	1.0	1.0 1.0
G FICL	1 40	-01 03	0.0	L D	211	1			T 40.0

Star	R.A. 1950	Decl. 1950	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
τ Pupp ε C Maj	h m 6 49 57	° ' -50 33 -28 54	2.8 1.6	G8 B1	,091 .005	" .025 .010	130 326	-0.2 -3.4	km. 'sec. +36.4* +27.4
ζ Gemi o ² C Maj δ C Maj μ Pupp η C Maj β C Min σ Pupp α ₁ Gemi α C Min	7 01 01 06 12 15 22 24 28 31 31 37	$\begin{array}{r} +20 & 39 \\ -23 & 45 \\ -26 & 19 \\ -44 & 33 \\ -37 & 00 \\ -29 & 12 \\ + & 8 & 23 \\ -43 & 12 \\ +32 & 00 \\ +32 & 00 \\ +5 & 21 \end{array}$	3.7-4.33.12.03.4-6.22.72.43.13.32.02.80.5	G0p B5p G4p M5e K5 B5p B8 M0 A2 A0 F5	.007 .006 .003 .332 .004 .007 .063 .191 .201 .209 1.242	.005 .007 .006 .018 .018 .012 .022 .016 .074 .074 .316	652 466 543 181 181 272 148 204 44 44	$\begin{array}{r} -2.8\\ -2.7\\ -4.1\\ -0.3\\ -1.0\\ -2.2\\ -0.2\\ -0.7\\ 1.4\\ 2.2\\ 3.0 \end{array}$	$\begin{array}{r} + \ 6.7^{*} \\ + 48.6 \\ + 34 \ 3^{*} \\ + 53 \ 0 \\ + 15.8 \\ + 40.4 \\ + 23 \ ^{*} \\ + 88.1^{*} \\ + \ 6.0^{*} \\ - \ 1.2^{*} \\ - \ 3.0^{*} \end{array}$
β Gemi ξ Pupp	42 47	$+28 09 \\ -24 44$	1.2 3.5	G9 K1	.623	. 105	31 543	$1.3 \\ -2.6$	+ 3.3 + 3.7*
ζ Pupp ρ Pupp γVelr ιε Cari υ Maj ιδ Velr ιε Hyda ζ Hyda ιε U Maj	8 02 05 08 21 26 43 44 53 56	$\begin{array}{r} -39 52 \\ -24 10 \\ -47 12 \\ -59 21 \\ +60 53 \\ -54 32 \\ + 6 36 \\ + 6 08 \\ +48 14 \end{array}$	2.3 2.9 2.2 1.7 3.5 2.0 3.5 3.3 3.1	08 F6 OW9 K0 G2 A0 F9 G7 A4	.032 .097 .002 .030 .166 .093 .193 .101 .500	.004 .025 .010 014 .030 012 .026 .060	815 130 326 233 109 272 125 54	$ \begin{array}{r} -4.7 \\ -0.1 \\ \\ -3.3 \\ -0.8 \\ -0.6 \\ -1.1 \\ 0.3 \\ 2.0 \\ \end{array} $	$\begin{array}{r} -24. \\ +46.6 \\ + 3.5 \\ +11.5 \\ +19.8 \\ + 2.2 \\ +36.8^{*} \\ +22.6 \\ +12.6 \end{array}$
λ Velr. β Cari. ι Cari. α Lync. κ Velr. α Hyda. θ U Maj. Ν Velr. ε Leon. ψ Cari.	9 06 13 16 18 21 25 30 30 43 46	$\begin{array}{r} -43 & 14 \\ -69 & 31 \\ -59 & 04 \\ +34 & 36 \\ -54 & 48 \\ -8 & 26 \\ +51 & 54 \\ -56 & 49 \\ +24 & 00 \\ -64 & 50 \end{array}$	$2.2 \\ 1.8 \\ 2.2 \\ 3.3 \\ 2.6 \\ 2.2 \\ 3.3 \\ 3.4-4.2 \\ 3.1 \\ 3.1$	K4 A0 F0 K8 B3 K4 F7 K5 G0 F0	.024 .192 .023 .214 .017 .036 1.096 .038 .045 .019	.016 022 .017 .018 .072 .022 .009	204 148 192 181 45 148 362 	-1.8 0 0 -1.2 -1.5 2.6 0.1 -2.1 	+18.4 -5. +13.3 +37.4 +21.7* -4.4 +15.8 -13.9 +5.1 +13.6
a Leon q Cari	10 06 15	$+12 13 \\ -61 05$	1.3 3.4	B6 K5	.244 .043	.046 .014	71 233	-0.4 -0.9	+ 2.6 + 8.6

Star		A. 1950	cl 1950			20.	vpe	11. Proper otion	ırallax	istance in ght Years	os. Mag.	ad. Vel.
		¥.	ן ב	í	2		T.	ΜĀ	L 9	EG	Ał	R
γ Leo	h . 10	m 17	。 +20	, 06	2.	3	G8	.347	.024	" 136	-0.8	km./sec -36.8
$\mu \cup Maj$	•	19	+41	45	3.	2	K4	.082	. 031	105	0.7	-20.3*
θ Cari	•	41	-64	08	3.	0	B0	.022	.007	466	-2.8	+24. *
η Cari	•	43	-59	25	1.0-	7.4	Pec	.007				-25.0
$\parallel \mu$ Velr	•	45	-49	09	2.	8	G5	.079	.033	99	0.4	+ 6.9
ν Hyda		47	-15	56	3.	3	K3	.218	020	163	-0.2	- 1.0
B U Maj	•	59	+56	39	2.	4	A3	.089	. 045	72	0.7	-12.1*
a U Maj	. 11	01	+62	01	2 .	0	G5	.137	036	91	-0.2	- 8.6*
$\psi \cup Maj$	·	07	+44	46	3.	2	K0	067	035	93	0.9	- 3.6
0 Leon	·	11	+20	47	2.	6	A2	208	058	56	1.4	-23.2
0 Leon	·	12	+15	42	3.	4	A2	. 103	025	130	04	+7.8
	·	33	-62	45	3.	3	Rð	.045	031	105	0.8	+ 7.9
p Leon	·	47	+14	51	2.	2	A2	. 507	084	39	1.8	- 2.3
γ U Maj	·	51	+53	58	2.	5	AO	.095	035	93	0.2	-11.1
δ Cent	. 12	06	-50	27	2 .	9	B3e	. 040	015	217	-1.2	+ 9 .
• Corv	.	08	-22	30	3.	2	K2	.063	024	136	0.1	+ 4.9
δ Cruc	.	12	-58	28	3.	1	B3	.045	017	192	-07	+26.4
δ U Maj	•	13	+57	19	3.	4	A0	.113	050	65	1.9	-12.
γ Corv	•	13	-17	16	2.	8	B8	.159	.024	136	-0.3	- 4.2*
a ¹ Cruc		24	-62	49	1.	6	B1	.048	022	148	-1.7	-12.2*
a ² Cruc		24	-62	49	2.	1	B3	.048	. 022	148	-1.2	+ 0.3*
$ \delta$ Corv	•	27	-16	14	3.	1	A0	.249	. 026	125	0.2	+ 8.7
γ Cruc		28	-56	50	1.	5	M4	.270		· · · · ·	••••	+21.3
β Corv	·	32	-23	07	2.	8	G5	.059	027	121	0.0	- 7.7
a Musc	·	34	-68	52	2 .	9	B5	.040	. 015	217	-1.2	+18.
$ \gamma$ Cent	·	39	-48	41	2.	4	A0	. 200	. 032	102	-0.1	- 7.5
$ \gamma $ Virg	·	39	- 1	10	2 .	9	F0	.561	080	41	2.4	-19.6
μρ Musc	·	43	-67	50	3.	3	B3	.039	.011	296	-1.5	+42.*
p Cruc	·	45	-59	25	1.	5	B1	.054	. 007	466	-4.3	-20 *
€ U Maj	·	52	+56	14	1.	7	A2	. 117	.067	49	0.8	-11.9*
a ² C. Ven		54	+38	35	2.	8	A1	.233	. 030	109	0.2	- 3.5
- X7.					_							
e Virg	13	00	+11	14	3	0	G6	.270	.037	88	0.8	-14.0
γ Hyda		16	-22	54	3.	3	G7	.085	.028	116	0.5	- 5.4
6 Cent		18	- 36	27	2.	9	A2	.351	.049	67	14	+ 0.1
IS* U. Maj		22	+55	11	2.	4	A2p	.131	.042	78	0.5	- 9.9*
a virg	·	23	-10	54	1.	2	B2	.051	.018	181	-2.5	+ 1.6*
y virg	1	32	- 0	20	3.	4	AZ	.285	.038	86	1.3	-13.1

Star		R.A. 1950	Decl. 1950		Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
 ε Cent. η U. Ma μ Cent ζ Cent. 		h m 13 37 46 47 52	$^{\circ}$ -53 +49 -42 -47	, 13 34 13 02	2.6 1.9 3.3 3.1	B2 B3 B3e B3	" .039 .116 .026 .080	" .012 .015 .009 .013	272 217 362 251	-2.0 -2.2 -1.9 -1.3	km./sec. - 5 6 - 10 9 + 12.6
η Boot. β Cent. π Hyda θ Cent. π Poot	•••••	52 14 00 04 04	+18 -60 -26 -36	39 08 26 07 26	2.8 0.9 3.5 2.3 0.2	GI B3 K3 G8 K0	.370 .039 .164 .745 2.287	. 100 . 026 . 037 . 056 102	33 125 88 58 32	-2.0 1.3 1.0 0.2	-12. * +27.2 + 1.3 - 5.1
γ Boot. γ Boot. η Cent. a Cent. a Circ. a Lupi	· · · · · · · · · · · · · · · · · · ·	30 32 36 38	+19 +38 -41 -60 -64 -46	20 32 56 38 46	3.0 2.6 0.1 3.4 2.9	A3 B3 G0 F0 B2	. 182 . 046 3. 682 . 308 033	.063 .012 .768 .063	52 272 4 52 362	$ \begin{array}{r} 0.2 \\ 2.0 \\ -2.0 \\ 4.5 \\ 2.4 \\ -2.3 \\ \end{array} $	$ \begin{array}{r} -35.5 \\ -0.2^{*} \\ -22.2^{*} \\ +7.4 \\ +7.3^{*} \end{array} $
$\begin{array}{c} \ \epsilon & \text{Boot.} \\ \ \epsilon & \text{Boot.} \\ \ a^* \text{ Libr.} \\ \beta & \text{U} & \text{Mi} \\ \beta & \text{Lupi.} \\ \kappa & \text{Cent.} \end{array}$		43 48 51 55 56	+27 -15 +74 -42 -41	17 47 22 56 54	2.7 2.9 2.2 2.8 3.4	G8 F1 K4 B3 B2	.045 .128 .028 .067 .034	.019 .056 .030 .012 .011	172 58 109 272 296	-0.9 1.6 -0.4 -1.8 -1.4	-16.4 -10. * +16 9 - 0.3* + 9.1*
	u	15 01 09 14 14	-25 -51 -68 -9	05 55 30 12	3.4 3.5 3.1 2.7	M4 G5 A0 B8	.091 .125 .064 .100	.020 .027 .015	163 121 217	-0.1 0.7 	- 4 3 - 9.7 0. -37.
δ Lupi γ U. Mi ι Drac. γ Lupi. α Cor. I	in	18 21 24 32 33	-40 +72 +59 -41 +26	28 01 08 00 53	3.4 3.1 3.5 3.0 2.3	B3 A2 K3 B3 A0	.031 .016 .010 .038 .160	.012 .022 .030 .013 .054	272 148 109 251 60	$ \begin{array}{c} -1.2 \\ -0.2 \\ 0.9 \\ -1.4 \\ 1.0 \end{array} $	+ 1.6 - 3.9* -11.1 + 6. + 1.0*
a Serp. β Tr. A π Scor. δ Scor.	u	42 51 50 57	2 + 6 - 63 - 25 - 22	35 17 58 29	2.8 3.0 3.0 2.5	K3 F0 B3 B1	.142 .436 .037 .039	.043 .096 .012 .011	76 34 272 296	$ \begin{array}{r} 1.0 \\ 2.9 \\ -1.6 \\ -2.3 \end{array} $	+ 3.0 - 0.3 - 3.0* -16. *
$\begin{array}{c c} \beta & \text{Scor.} \\ \delta & \text{Ophi.} \\ \epsilon & \text{Ophi.} \\ \sigma & \text{Scor.} \\ \eta & \text{Drac.} \end{array}$		16 03 19 10 18 23	$\begin{vmatrix} -19 \\ -3 \\ -4 \\ -25 \\ +61 \end{vmatrix}$	40 34 34 28 38	2.8 3.3 3.3 3.1 2.9	B3 K8 G9 B1 G5	.029 .159 .088 .033 .062	.016 .030 .031 .009 .038	204 109 105 362 86	$ \begin{bmatrix} -1.2 \\ 0.7 \\ 0.8 \\ -2.1 \\ 0.8 \end{bmatrix} $	$ \begin{vmatrix} -9.3^{*} \\ -19.8 \\ -10.3 \\ -0.4^{*} \\ -14.3 \end{vmatrix} $

Star	R.A. 1950	Decl. 1950	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
	l h m	.0	1	1			1	1	11
									km./sec
la Scor	16 26	-26 19	1.2	M1	.032	.019	172	-2.4	- 3.2*
β Herc	28	+21 36	2.8	G4	.104	.020	163	-0.7	-25.8^{*}
τ Scor	33	-2807	2.9	B1	.037	.009	362	-2.3	+ 0.6
ζ Ophi	34	-1028	2.7	BO	.023	.008	407	-2.8	-19 *
Il Herc.	39	+31 42	3.0	GO	601	105	31	3 1	-70.8*
a Tr. Au	43	-68 56	1 0	K5	031	025	130	_1 1	- 3 7
6 Scor	47	-24 12	21.0		665	020	100	-1.1	- 3.7
ul Scor	10	97 50	2.1	D2.	.000	011	00	0.3	- 2.5
μ. Οτοι	40	-37 38	0.1	Бар	.030	.011	290	-1.7	
(Arae	54	-55 55	3.1	K5	.046	.028	116	0.3	- 6.0
k Opni	55	+ 9 27	3.1-4.0	K3	.290	.042	78	1.2	-55.6
lln Ophi	17 00	15 40	0.6	4.9	005	047	60	1.0	1.0
	11 00	-15 40	2.0	AZ	.095	.047	09	1.0	- 1.0
η Scor	08	-43 11	3.4	A7	.294	.066	49	2.5	-28.4
§ Drac	09	+65 47	3.2	B8	.023	028	116	0.4	-14.1
a' Herc	12	+14 27	3.1-3.9	M7	.030	.008	407	-2.4	-32.5
ð Herc	13	+2454	3.2	A2	.164	.036	91	1.0	-39. *
π Herc	13	+3652	3.4	K3	.021	.018	181	-0.3	-25.7
θ Ophi	19	-24 57	3.4	B2	.031	.008	407	-2.1	- 3.6
β Arae	21	-55 29	2.8	K1	.036	.023	142	-0.4	- 0.4
v Scor	27	-37 15	2.8	B3	.042	.010	326	-2.2	+18. *
a Arae	28	-49 50	3.0	B3e	.090	.015	217	-1.1	- 2.2
B Drac	29	$+52\ 20$	3.0	GO	012	007	466	-2.8.	-20.1
λ Scor	30	-37 04	17	B2	036	016	204	-2.3	0 +
a Ophi	33	± 12.35	21	Δ <u>Ω</u>	264	060	54	10	U. 115 •
A Scor	24	19 50	2.1	FO	.204	.000	196	1.0	+10.
# Scor	20	-42 00	2.0		.012	.024	130	-1.1	+ 1.4
R Och:	39 41	-39 00	2.5	Bo	.028	.009	362	-2.7	-10. •
ρ Opm	41	+ 4 35	2.9	K2	.157	.030	109	0.3	-11.9
Scor	44	-40 06	3.1	F8	.004	.008	407	-2.4	-27.6*
μ Herc	44	+27 45	3.5	G5	.817	114	28	3.8	-16.1
G Scor	46	-37 02	3.2	K2	.069	. 029	112	0.5	+24.7
v Ophi	56	- 946	3.5	G7	.118	.022	148	0.2	+12.4
γ Drac	55	+51 30	2.4	K5	.026	.026	125	-0.5	-27.8
γ Sgtr	18 03	-30 26	3.1	K0	. 202	. 030	109	0.5	+22.3*
η Sgtr	14	-36 47	3.2	M4	.216	.030	109	0.6	+ 0.5
δ Sgtr	18	-29 51	2.8	K4	.052	.033	99	0.4	-20.0
η Serp	19	- 2 55	3.4	G9	. 898	.050	65	1.9	+ 8.9
é Sgtr	21	-34 25	2.0	A0	.139	.020	163	-1.5	-10.8
λ Sgtr	25	-25 27	2.9	К1	.196	.036	91	0.7	-43.3
a Lyra	35	+38 44	0.1	AI	.348	.140	23	0.8	-13.8
	!							0.01	10.0
Star	R.A. 1950	Decl. 1950	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
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φ Sgtr β Lyra σ Sgtr γ Lyra Sgtr	h m 18 43 48 52 57 59	$\begin{array}{c} \circ & -27 & 03 \\ +33 & 18 \\ -26 & 22 \\ +32 & 37 \\ -29 & 57 \end{array}$	3.33.4-4.12.13.32.7	B8 B2p B3 B9p A2	.150 .011 .067 .008 .019	" .015 .006 .021 .016 .035	217 543 155 204 93	-0.8 -2.7 -1.3 -0.7 0.4	km./sec. +21.5* -19.0* -10.7 -21.5* +22.1
ζ Aqil τ Sgtr δ Drac δ Aqil β' Cygn γ Aqil α Aqil	19 03 04 07 13 23 29 43 44 48	$\begin{array}{r} +13 & 47 \\ -27 & 45 \\ -21 & 06 \\ +67 & 34 \\ + & 3 & 01 \\ +27 & 51 \\ +45 & 00 \\ +10 & 29 \\ + & 8 & 44 \end{array}$	3.0 3.4 3.0 3.2 3.4 3.2 3.0 2.8 0.9	A0 K0 F2 G8 A3 K0 A1 K3 A2	. 103 . 268 . 041 . 135 . 267 . 010 . 067 . 018 . 659	.038 .036 .017 .028 .052 .010 .023 .018 .184	86 91 192 116 63 326 116 181 18	$\begin{array}{c} 0.9 \\ 1.2 \\ -0.8 \\ 0.4 \\ 2.0 \\ -1.8 \\ 0.2 \\ -0.9 \\ 2.2 \end{array}$	$\begin{array}{r} -25. * \\ +45.4* \\ -9.8 \\ +24.8 \\ -32.3* \\ -23.9* \\ -20. \\ -2.0 \\ -26.1 \end{array}$
 θ Aqil β Capr γ Cygn a Pavo a Indi a Cygn ϵ Cygn 	20 09 18 20 22 34 40 44	$\begin{array}{rrrrr} - & 0 & 58 \\ - & 14 & 56 \\ + & 40 & 06 \\ - & 56 & 54 \\ - & 47 & 28 \\ + & 45 & 06 \\ + & 33 & 47 \end{array}$	3.43.22.32.13.21.32.6	A0 F8 F8 B3 G2 A2p G7	.035 .042 .006 .087 .072 .004 .485	.018 .022 .008 .014 .034 .002 .040	181 148 407 233 96 1630 81	$ \begin{array}{r} -0.3 \\ -0.1 \\ -3.2 \\ -2.2 \\ 0.9 \\ -7.2 \\ 0.6 \\ \end{array} $	$\begin{array}{r} -28.6^{*} \\ -19.0^{*} \\ -7.6 \\ +1.8^{*} \\ -1.1 \\ -6.3^{*} \\ -10.5^{*} \end{array}$
ζ Cygn a Ceph β Ceph β Aqar ε Pegs δ Capr γ Grus	21 11 17 28 29 42 44 51	$\begin{array}{r} +30 \ 01 \\ +62 \ 22 \\ +70 \ 20 \\ -5 \ 48 \\ +9 \ 39 \\ -16 \ 21 \\ -37 \ 36 \end{array}$	3.42.63.3-3.43.12.53.03.2	G6 A2 B1 G1 K2 A3 B8	.061 .163 .013 .020 .028 .395 .114	.018 .076 .006 .008 .014 .062 .020	181 43 543 407 233 53 163	$ \begin{array}{r} -0.3 \\ 2.0 \\ -2.8 \\ -2.4 \\ -1.8 \\ 2.0 \\ -0.3 \end{array} $	$+16.9^*$ - 8. - 7.2 + 6.7 + 5.2 - 6.4^* - 2.1
a Aqar a Grus a Tucn β Grus η Pegs a Psc. A	22 03 05 15 40 41 55	$ \begin{array}{r} - & 0 & 34 \\ -47 & 12 \\ -60 & 31 \\ -47 & 09 \\ +29 & 58 \\ -29 & 53 \\ \end{array} $	3.2 2.2 2.9 2.2 3.1 1.3	G0 B5 K5 M6 G1 A3	.019 .202 .088 .131 .039 .367	.006 .036 .019 .010 .016 .118	543 91 172 326 204 28	$ \begin{array}{c} -2.9 \\ 0.0 \\ -0.7 \\ -2.8 \\ -0.9 \\ 1.7 \\ 0.6 \\ \end{array} $	+ 7.6 +11.8 +42.2* + 1.6 + 4.4* + 6.5
 β Pegs a Pegs γ Ceph 	23 01 02 37	$ +27 49 \\ +14 56 \\ +77 21$	2.6 2.6 3.4	M3 A0 K1	.235 .077 .167	.020 .033 .062	163 99 53	-0.9 0.2 2.4	+ 8.6 - 4.* -42.0

The star clusters for this observing list have been selected to include the more conspicuous members of the two main classes—open clusters and globular clusters. Most of the data are from Shapley's Star Clusters and from Trumpler's catalogue in Lick Bulletin No. 420. In the following table N.G.C. indicates the serial number of the cluster in the New General Catalogue of Clusters and Nebulae; M, its number in Messier's catalogue; Con, the constellation in which it is located; a and δ , its right ascension and declination; Cl, the kind of cluster, Op for open or galactic and Gl for globular; Diam., the apparent diameter in minutes of arc; Mag. B.S., the magnitude of the fifth brightest star in the case of open clusters, the mean of the 25 brightest for globular; No., the number of stars in the open clusters down to the limiting magnitudes of the photographs on which the particular clusters; and Dist, the distance in light years.

N.G.C.	M	Con.	a	19	50	δ	C1.	Diam.	Mag.	No.	Int.	Dist.
			hr	n	•	'		,	B.S.		mag.	l.y
869		h Per	02 15	5.5	+56	55	Op	30	7			4,300
884		χPer	02 18	3.9	+56	53	Op	30	7			4,300
1039	34	Per	02 38	3.3	+42	35	Op	30	9	80		1,500
Pleiades	45	Tau	03 44	1.5	+23	58	Op	120	4.2	250		490
Hyades		Tau	04 17	'	+15	30	Op	400	4.0	100		1 20
1912	38	Aur	05 25	5.3	+35	48	Op	18	9.7	100		2,800
2 099	37	Aur	05 49).0	+32	33	Op	24	9.7	150		2,700
2 168	35	Gem	06 05	5.7	+24	21	Op	29	9.0	120		2,700
2287	41	C Ma	06 44	.9	-20	42	Op	32	9	50		1,300
2632	44	Cnc	08 37	7.2	+20	10	Op	90	6.5	350		490
5139		ωCen	13 23	3.7	-47	03	Gl	23	12.9		3	22,000
5272	3	C Vn	13 39	9.9	+28	38	Gl	10	14.2		4.5	40,000
5904	5	Ser	15 15	5.9	+02	16	Gl	13	14.0		3.6	35,000
6121	4	Scr	16 20).5	-26	24	Gl	14	13.9		5.2	24,000
6205	13	Her	16 39	9.9	+36	33	Gl	10	13.8		4.0	34,000
6218	12	Oph	16 44	.6	-01	51	GI	9	14.0		6.0	36,000
6254	10	Oph	16 54	1.5	-04	02	Gl	8	14.1		5.4	36,000
6341	92	Her	17 15	5.6	+43	12	Gl	8	13.9		5.1	36,000
6494	23	Sgr	17 54	l.0	-19	01	O p	27	10.2	120		2,200
6611	16	Ser	18 16	5.0	-13	48	Op	8	10.6	55		6,700
6656	22	Sgr	18 33	3.3	-23	57	GI	17	12.9		3.6	2 2, 000
7078	15	Peg	21 27	.6	+11	57	Gl	7	14.3		52	4 3 ,Cu0
7089	2	Aqr	21 30).9	-01	04	Gl	8	14.6		5.0	45,000
7092	39	Cyg	21 30).5	+48	13	Op	32	6.5	25		1,000
7654	52	Cas	23 22	2.0	+61	19	Op	13	11.0	120		4,400

GALACTIC NEBULAE

The galactic nebulae here listed have been selected to include the most readily observable representatives of planetary nebulae such as the Ring Nebula in Lyra, diffuse bright nebulae like the Orion nebula and dark absorbing nebulosities such as the Coal Sack. These objects are all located in our own galactic system. The first five columns give the identification and position as in the table of clusters. In the Cl column is given the classification of the nebula, planetary nebulae being listed as Pl, diffuse nebulae as Dif, and dark nebulae as Drk. Size indicates approximately the greatest apparent diameter in minutes of arc; and m n is the magnitude of the planetary nebula and m * is the magnitude of its central star. The distance is given in light years, and the name of the nebulae is added for the better known objects.

N .G.C.	М	Con	h	a 19 m	950 δ	,	CI	Size	m n	m *	Dist. 1.y.	Name
650	76	Per	01	38.3	+51	20	Pl	1.5	11	17	15,000	
1952	1	Tau	05	31.5	+21	59	Pl	6	11	16	10,000	Crab
1976	42	Ori	05	32.5	-05	25	Dif	30			1,800	Orion
B 33		Ori	05	38.0	-02	29	Drk	4			300	Horsehead
2261		Mon	06	36.4	+08	47	Dif	2				Hubble's var
2392		Gem	07	26.2	+21	02	Pl	0.3	8	10	2.800	
2440		Pup	07	39.6	-18	05	Pl	0.9	11	16	8.600	
3587	97	UMa	11	11.8	+55	17	P1	3.3	11	14	12,000	Owl
		Cru	12	48	-63		Drk	300			300	Coalsack
6210		Her	16	42 .4	+23	54	P1	0.3	10	12	5,600	
D 79		Only	17	90 F		26	Dala	20			400	C
B12 6514	00	Opn	17	20.0	-23	30 00	Drk	20			400	S nebula
0014 D06	20	Sgr	17	59.5	-23	02		24			3,200	Irind
D80 6592	0	Sgr	11	00.6	-21	02 92	Drk	50 50			2 600	Lawren
0020 6543	0	Dro	10	58.6	- 24	20 20	Dil	0.4	0	11	3,000	Lagoon
0040		Dia	11	30.0	700	90	11	0.4	9	11	3,500	
6572		Oph	18	10.2	+06	50	Pl	0.2	9	12	4,000	
B92		Sgr	18	12.7	-18	15	Drk	15				
6618	17	Sgr	18	18.0	-16	12	Dif	26			3,000	Horseshoe
6720	57	Lyr	18	5 2.0	+32	58	P1	1.4	9	14	5,400	Ring
6826		Cyg	19	43.5	+50	2 4	Pl	0.4	9	11	3,400	
6853	27	Vul	19	57.4	+22	35	P1	8	8	13	3.400	Dumb-bell
6960		Cyg	20	43.6	+30	32	Dif	60	-		-,	Network
7000		Cyg	20	57.0	-+44	07	Dif	100				N. America
7009		Agr	21	01.4	-11	34	Pl	0.5	8	12	3,000	
7662		And	23	23.4	+42	12	Pl	0.3	9	13	3,900	

EXTRA-GALACTIC NEBULAE

Among the hundreds of thousands of systems far beyond our own galaxy relatively few are readily seen in small telescopes. The following list contains a selection of the closer brighter objects of this kind. The first five columns give the catalogue numbers, constellation and position on the celestial sphere. In the column Cl, E indicates an elliptical nebula, I an irregular object, and Sa, Sb, Sc spiral nebulae, in which the spiral arms become increasingly dominant compared with the nucleus as we pass from a to c. The remaining columns give the apparent magnitude of the nebula, its distance in light years and the radial velocity in kilometers per second. As these objects have been selected on the basis of ease of observation, the faint, very distant objects which have spectacularly large red shifts, corresponding to large velocities of recession, are not included.

N.G.C.	М	Con	a 19	50 δ	Cl	Dimens.	Mag.	Distance	Vel.
			hm	• /		, ,		1.y.	km/sec
221	32	And	00 39.9	+40 36	Е	3×3	8.8	800,000	- 185
224	31	And	00 40.0	+41 00	Sb	160×40	5.0	800,000	- 220
SMC		Tuc	00 53	$-72\ 38$	Ι	220×220	1.5	100,000	+ 170
598	33	Tri	01 31.0	+3024	Sc	60×40	7.0	700,000	- 70
LMC		Dor	05 21	$-69 \ 27$	I	430×530	0.5	90,000	+ 280
3031	81	UMa	09 51.5	+69 18	Sb	16×10	8.3	2,400,000	- 30
3034	82	UMa	09 51.8	+6958	Ι	7×2	9.0	2,600,000	+ 290
3368	96	Leo	10 44.1	+12 05	Sa	7×4	10.0	5,700,000	+ 940
3623	65	Leo	11 16.3	+13 22	Sb	8×2	9.9	5,000,000	+ 800
3627	66	Leo	11 17.6	+13 16	Sb	8× 2	9.1	4,300,000	+ 650
4258		CVn	12 16.5	+47 34	Sb	20×6	8.7	4,600,000	+ 500
4374	84	Vir	$12 \ 22.5$	+13 09	E	3×2	9.9	6,000,000	+1050
4382	85	Com	$12 \ 22.9$	+18 28	Е	4×2	10.0	3,700,000	+ 500
4472	49	Vir	12 27. 2	+08 16	E	5×4	10.1	5,700,000	+ 850
456 5		Com	12 33.9	$+26\ 16$	Sb	15×1	11.0	7,600,000	+1100
4594		Vir	12 37.4	-11 20	Sa	7× 2	9.2	7,200,000	+1140
4649	60	Vir	12 41.1	+11 50	E	4×3	9.5	7,500,000	+1090
4736	94	CVn	$12 \ 48.6$	+41 24	Sb	5×4	8.4	3,00 0,000	+ 290
4826	64	Com	12 54.3	+21 57	Sb	8×4	9.2	1,300,000	+ 150
5005		CVn	13 08.6	+37 20	Sc	5×2	11.1	6,600,000	+ 9 0 0
5055	63	CVn	13 13.6	+42 18	\mathbf{Sb}	8× 3	9.6	3,600,000	+ 450
5194	51	CVn	13 27.8	+47 27	Sc	12×6	7.4	3,000,000	+ 250
5236	83	Hya	$13 \ 34.2$	-29 36	Sc	10×8	8	2,900,000	+ 500
6822		Sgr	$19 \ 42.4$	-1453	1	20×10	11	1,000,000	- 150
7331		Peg	22 34.8	+33 59	Sb	9× 2	10.4	5,200,000	+ 500



Mi	idnig	ht	t.		•	• •	•	 •	.Feb.	6
11	p.m.	•						 •	. "	21
10	**								. Mar.	7
9	* 4			•					. "	22
8	"			•				 •	Apr.	6
7	"			•				 •	"	21

The centre of the map is the zenith, the circumference the horizon. To identify the stars hold the map so that the part of the horizon you are facing is down.



M:	id nig	h	t.	•	•	•	•	•	•	•	•	•	. May	8
11	p.m.									•			. "	24
10	**								•				. June	7
9	"												. "	22
8	"								•	•	•	•	. July	6

The centre of the map is the zenith, the circumference the horizon. To identify the stars hold the map so that the part of the horizon you are facing is down.



Mi	id nig l	ht	• •	 •	 	•	Aug.	5
11	p.m.				 		44	21
10	- 44				 		Sept.	7
9	"	• •			 			23
8	"				 	•	Oct.	10
7	**				 	•	. "	26
6	**				 	•	Nov.	6
5	"		•		 		. "	21

The centre of the map is the zenith, the circumference the horizon. To identify the stars hold the map so that the part of the horizon you are facing is down.



0	
11 p.m " 2	21
10 " Dec.	6
9 " " 2	21
8 "Jan.	5
7 " " 2	20
6 " Fe b.	6

The centre of the map is the zenith, the circumference the horizon. To identify the stars hold the map so that the part of the horizon you are facing is down.

EPHEMERIS FOR THE PHYSICAL OBSERVATION OF THE SUN

Dat	e	Р	B°	L。	Date	Р	В。	L.
·		o	 o	0			o	•
Jan. Feb.	$1 \\ 6 \\ 11 \\ 16 \\ 21 \\ 26 \\ 31 \\ 5 \\ 5 \\ 10 \\ 10 \\ 10 \\ 10 \\ 10 \\ 10 $	$\begin{array}{r} + 2.39 \\ - 0.04 \\ - 2.46 \\ - 4.84 \\ - 7.15 \\ - 9.38 \\ - 11.52 \\ - 13.55 \end{array}$	$\begin{array}{r} -3.02 \\ -3.59 \\ -4.14 \\ -4.65 \\ -5.12 \\ -5.56 \\ -5.95 \\ -6.29 \end{array}$	$\begin{array}{r} 66.78 \\ 0.93 \\ 295.09 \\ 229.25 \\ 163.41 \\ 97.58 \\ 31.74 \\ 325.92 \end{array}$	Jul. 5 10 15 20 25 30 Aug. 4 9	$\begin{array}{r} -1.23 \\ +1.04 \\ +3.29 \\ +5.49 \\ +7.65 \\ +9.73 \\ +11.73 \\ +13.64 \end{array}$	$\begin{array}{r} +3.30 \\ +3.82 \\ +4.32 \\ +4.79 \\ +5.22 \\ +5.63 \\ +5.99 \\ +6.31 \end{array}$	$\begin{array}{c} 145.07\\78.90\\12.73\\306.57\\240.42\\174.28\\108.15\\42.04\end{array}$
Mar	10 15 20 25 2	-15.45 -17.23 -18.86 -20.35 -21.68	$ \begin{array}{r} -6.59 \\ -6.83 \\ -7.02 \\ -7.17 \\ -7.24 \end{array} $	$\begin{array}{r} 260.08 \\ 194.25 \\ 128.40 \\ 62.55 \\ 356.60 \end{array}$	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	+15.44 +17.13 +18.70 +20.14 +21.45	+6.59 +6.82 +7.00 +7.13 +7.22	$\begin{array}{c} 335.93 \\ 269.84 \\ 203.77 \\ 137.71 \\ 71.66 \end{array}$
11141.	7 12 17 22	-21.03 -22.86 -23.87 -24.72 -25.39 25.00	-7.24 -7.25 -7.22 -7.13 -6.99 6.70	$\begin{array}{r} 330.09\\ 290.82\\ 224.94\\ 159.04\\ 93.12\\ 27.10\end{array}$	8 13 18 23	+21.33 +22.62 +23.64 +24.51 +25.22 +25.76	+7.22 +7.25 +7.23 +7.16 +7.03	$\begin{array}{r} 5.62 \\ 299.60 \\ 233.59 \\ 167.59 \\ 101.60 \end{array}$
Apr.	1 6 11 16 21 26	$\begin{array}{r} -25.90 \\ -26.22 \\ -26.37 \\ -26.33 \\ -26.11 \\ -25.71 \\ 95.12 \end{array}$	$ \begin{array}{r} -6.79 \\ -6.55 \\ -6.25 \\ -5.92 \\ -5.54 \\ -5.12 \\ 4.66 \\ \end{array} $	$\begin{array}{r} 27.19\\ 321.23\\ 255.26\\ 189.27\\ 123.26\\ 57.23\\ 251.17\end{array}$	Oct 3 8 13 18 23 28	+25.76 +26.14 +26.34 +26.36 +26.19 +25.84	+0.83 +6.63 +6.35 +6.02 +5.65 +5.24	$\begin{array}{c} 101.00\\ 35.62\\ 329.65\\ 263.69\\ 197.74\\ 131.79\\ 65.85\end{array}$
May	20 1 6 11 16 21 26	$\begin{array}{r} -23.13 \\ -24.36 \\ -23.40 \\ -22.28 \\ -20.98 \\ -19.51 \\ -17.90 \end{array}$	$\begin{array}{r} -4.00 \\ -4.18 \\ -3.66 \\ -3.12 \\ -2.56 \\ -1.98 \\ -1.40 \end{array}$	$\begin{array}{c} 351.17\\ 285.10\\ 219.01\\ 152.91\\ 86.78\\ 20.64\\ 314.49\end{array}$	Nov. 2 7 12 17 22 27	$\begin{array}{r} +23.29\\ +24.54\\ +23.59\\ +22.46\\ +21.12\\ +196.0\\ +179.1\end{array}$	$\begin{array}{r} +4.78 \\ +4.29 \\ +3.76 \\ +3.20 \\ +2.62 \\ +2.02 \\ +1.40 \end{array}$	$\begin{array}{c} 03.83\\ 359.92\\ 293.99\\ 228.07\\ 162.16\\ 96.25\\ 30.35\end{array}$
Jun.	31 5 10 15 20 25 30	$\begin{array}{r} -16.14 \\ -14.25 \\ -12.25 \\ -10.16 \\ -7.98 \\ -5.77 \\ -3.51 \end{array}$	$\begin{array}{c} -0.80 \\ -0.19 \\ +0.41 \\ +1.01 \\ +1.60 \\ +2.18 \\ +2.75 \end{array}$	$\begin{array}{c} 248.34 \\ 182.16 \\ 115.99 \\ 49.81 \\ 343.62 \\ 277.43 \\ 211.25 \end{array}$	De 2 7 12 17 22 27 Jan. 1	$\begin{array}{c} +16.05 \\ +14.04 \\ +11.90 \\ + 9.66 \\ + 7.32 \\ + 4.93 \\ + 2.50 \end{array}$	$\begin{array}{ c c c c c } +0.77 \\ +0.13 \\ -0.51 \\ -1.15 \\ -1.78 \\ -2.39 \\ -2.99 \end{array}$	$\begin{array}{c c} 324.45\\ 258.56\\ 192.68\\ 126.80\\ 60.93\\ 355.07\\ 289\ 22\end{array}$

P—The position angle of the axis of rotation, measured eastward from the north point of the disk.
B_o—The heliographic latitude of the centre of the disk.
L_o—The heliographic longitude of the centre of the disk, from Carrington's solar meridian.

Carrington's Rotation Numbers-Greenwich date of commencement of synodic rotations.

No.	Commences	No.	Commences	No.	Commences
1301	1950 Dec. 9.74	1306	1951 Apr. 25.33	1311	1951 Sep. 8.43
1 302	1951 Jan 6.07	1307	May 22.56	1312	Oct. 5.70
1303	Feb. 2.41	1308	Jun. 18.76	1313	Nov. 1.99
1304	Mar. 1.75	1309	Jul. 15.96	1314	Nov. 29.30
1305	Mar. 29.06	1310	Aug. 12.18	1315	Dec. 26.63

Continued from page 57.

all other features observed. Information equally important, but often forgotten, is the exact time and date of the phenomenon and an accurate description of where the observer was situated, given within 100 yds. if possible.

Skilled visual or photographic observations from two or more stations make possible the computation of meteor heights. Most meteors are visible in the range from 40 to 80 miles above the earth's surface and move with velocities ranging from 20 to 60 miles per second.

METEORS AND METEORITES

Many common terrestrial stones have mistakenly been thought to have a meteoric origin, and any supposed meteorite should be investigated carefully. Contrary to popular belief, meteorites do not contain valuable minerals in quantities sufficient to make them of commercial interest, but they have a definite scientific value. Meteorites are of two main types, iron and stone. The irons have specific gravity ranging from 7 to 8 and are amost entirely composed of metallic nickel-iron. The stones have a specific gravity ranging from 2 to 4 or greater and, with very few exceptions, contain metallic inclusions that are revealed on grinding or filing the specimen. A freshly fallen meteorite is covered by a smooth black fusion crust but oxidation removes this where the object has lain in the ground for any length of time. Any object whose history and structure indicate that it is of meteoric origin should be submitted to some authority for further study.

A more detailed discussion of both visual and photographic observations of meteors will be found in "General Instructions for Meteor Observing." Meteor observations for the United States may be sent to the American Meteor Society, Flower Observatory, Upper Darby, Pa.; those for Canada to the writer at the Dominion Observatory, Ottawa, Ont.

Shower	Approx a	. Radiant δ	Current Maximum Date	Spectacular Displays	Hourly Number (all meteors)	Duration (in days)	Abbre- viations (for use in observing records)
Quadrantids	232°	+52°	Jan. 3		20	4	Q
Lyrids	280	+37	Apr. 21		10	4	Ŷ
Eta Aquarids	336	- 1	May 4		10	8	Е
Delta Aquarids .	340	-17	July 28		20	12	D
Perseids	47	+57	Aug. 12		50	25	Р
Giacobinids	267	+55	Oct. 9	1933, 1946		1	J
Orionids	96	+15	Oct. 22		20	14	0
Taurids	56	+16	Nov. 10?			30	Т
Leonids	152	+22	Nov. 16	1799, 1833,	20	14	L
				1866, 1867			
Bielids	25	+45	Nov. 27	1872, 1885			В
Geminids	110	+33	Dec. 12		30	14	G

PRINCIPAL METEOR SHOWERS FOR THE NORTHERN HEMISPHERE

TABLE OF PRECESSION FOR 50 YEARS

-	R.A.	h m	12 00	11 30	11 00	10 30	10 00	9 30	00 6	8 30	8 00	7 30	2 00	6 30	6 00	00 00	03 30	23 00	00 00	00 22	22 00	21 30	21 00	20 30	20 00	19 30	19 00	18 30
Prec.	Dec.		- 16.7	- 16.6	- 16.1	- 15.4	- 14.5	- 13.2	- 11.8	- 10.2	- 8.3	- 6.4	- 4.3	- 2.2	0.0		1 18 8	1 18 1	T-01 1	+ 10.4	+ 14.5	+13.2	+ 11.8	+10.2	+ 8.3	+ 6.4	+ 4.3	4 2.9
	-30°	E	+2.56	2.48	2.39	2.31	2.24	2.17	2.11	2.05	2.00	1.97	1.94	1.92	1.92	0 1 0	9 84	2.73		10.2	2.88	2.95	3.02	3.07	3.12	3.16	3.18	3.20
	-20°	B	+2.56	2.51	2.45	2.40	2.36	2.31	2.27	2.24	2.21	2.19	2.17	2.16	2.16	0 - 0	9.61	2.87	0	21.2	2.76	2.81	2.85	2.88	2.91	2.93	2.95	9.06
	-10°	E	+2.56	2.53	2.51	2.49	2.46	2.44	2.42	2.40	2.39	2.38	2.37	2.37	2.36	1 0 10	9 50	9.61	10.0	2.04	2.66	2.68	2.70	2.72	2.73	2.74	2.75	9.75
	00	Ш	+2.56	2.56	2.56	2.56	2.56	2.56	2.56	2.56	2.56	2.56	2.56	2.56	2.56	0 0 .	9 56	9.56	00.0	00.2	2.56	2.56	2.56	2.56	2.56	2.56	2.56	9.56
	+10°	В	+2.56	2.59	2.61	2.64	2.66	2.68	2.70	2.72	2.73	2.74	2.75	2.75	276	1 0 10	7 2.00	9.51	10.40	2.43	2.46	2.44	2.42	2.40	2.39	2.38	2.37	2.37
ension	+20°	В	+2.56	2.61	2.67	2.72	2.76	2.81	2.85	2.88	2.91	2.93	2.95	2.96	2.97	0 1 0 1	9.51	2.45	01.0	2.40	2.36	2.31	2.27	2.24	2.21	2.19	2.17	2.16
ght Asc	+30°	8	+ 2.56	2.64	2.73	2.81	2.88	2.95	3.02	3.07	3.12	3.16	3.18	3.20	3.20	0101	7 2.00	03.30		10.2	2.24	2.17	2.11	2.05	2.00	1.97	1.94	1.92
n in Ri	+40°	B	+2.56	2.68	2.80	2.92	3.03	3.13	3.22	3.30	3.37	3.42	3.46	3.49	3.50	0 10	2.44	9.39		2.20	2.09	1.99	1.90	1.81	1.75	1.70	1.66	1.63
recessio	+50°	B	+2.56	2.73	2.90	3.07	3.22	3.37	3.50	3.61	3.71	3.79	3.84	3.88	3.89	0 - 0 -	00.2 1	66.6		00.2	1.90	1.75	1.62	1.51	1.41	1.33	1.28	1 25
P	+60°	m	+ 2.56	2.81	3.06	3.30	3.52	3.73	3.92	4.09	4.23	4.34	4.42	4.47	4.49	0 0 1	9.31	2.06	00.	70.1	1.60	1.39	1.20	1.03	+0.89	+ 0 78	+ 0.70	+ 0.65
	+70°	E	+2.56	2.96	3.36	3.73	4.09	4.42	4.73	4.99	5.21	5.39	5.52	5.60	5.62	0 4 0 1	9.16	77 1		RC.I	1.03	0.70	+ 0.40	+0.13	- 0.09	- 0.27	- 0.40	- 0.47
	+75°	B	+2.56	3.10	3.64	4.15	4.64	5.09	5.50	5.86	6.16	6.40	6.58	6.68	6.72	0 0 1	00.2 1	1 48		18.0	+0.46	+0.03	- 0.38	- 0.74	- 1.04	- 1.28	- 1.45	- 1.56
	+80°	в	+2.56	3.38	4.19	4.98	5.72	6.40	7.02	7.57	8.03	8.40	8.66	8.82	8.88	01 0 1	1 89	4 0 03	1 0 1	+ 0.14	- 0.60	- 1.28	- 1.90	- 2.45	- 2.91	- 3.27	- 3.54	- 3.70
	= +85°	E	+ 2.56	+ 4.22	+ 5.85	+ 7.43	+ 8.92	+10.31	+11.56	+12.66	+13.58	+14.32	+14.85	+15.18	+15.29	0 A 0 A	1 0.00	- 0.73	22.0	- 2.31	- 3.80	- 5.19	- 6.44	- 7.54	- 8.46	- 9.20	- 9.73	-10.08
Prec.	Dec. 8		+ 16.7	+ 16.6	+ 16.1	15.4	+ 14.5	+ 13.2	+ 11.8	+ 10.2	+ 8.3	+ 6.4	+ 4.3	+ 2.2	+ 0.0		- 18 B	18.1	1.01	10.4	- 14.5	- 13.2	- 11.8	- 10.2	- 8.3	- 6.4	- 4.3	0.0 -
	R.A.	h m	0 00 0	0 30	1 00	1 30	- 00 -	2 30	3 00	3 30	4 00	4 30	5 00	5 30	6 00	00 01	10 20	12 00	00 01	13 30	14 00	14 30	15 00	15 30	16 00	16 30	17 00	17 30

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